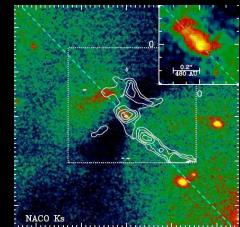
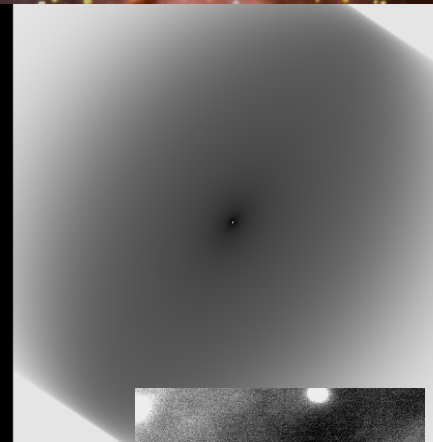
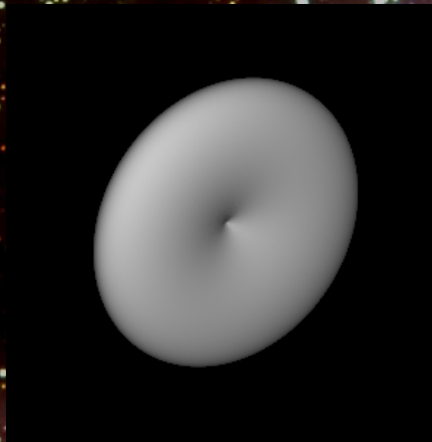
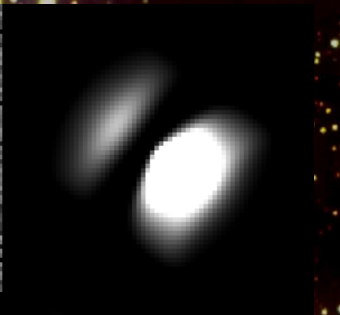
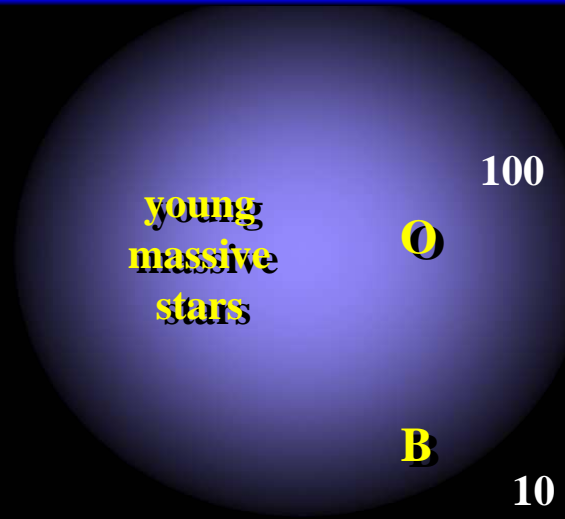


Evidence for Disks around YMSs from 3D Radiative Transfer Modeling

Jürgen Steinacker

Max-Planck-Institute for Astronomy Heidelberg (MPIA)
Institute for Computational Astronomy
at the Center for Astronomy Heidelberg (ZAH)





young massive stars

O

B

Herbig Ae/Be stars

A

T Tauri stars

F

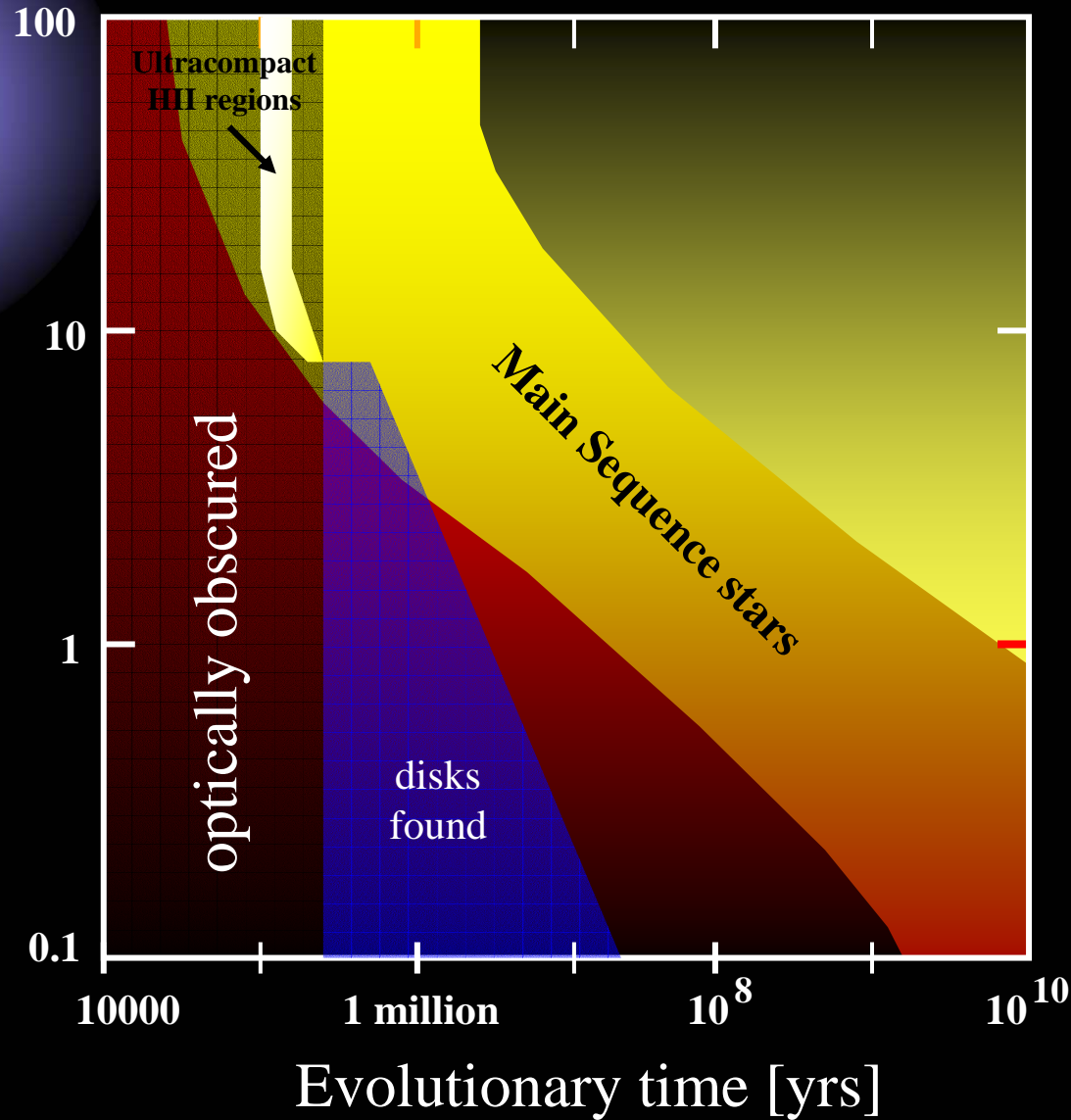
G

K

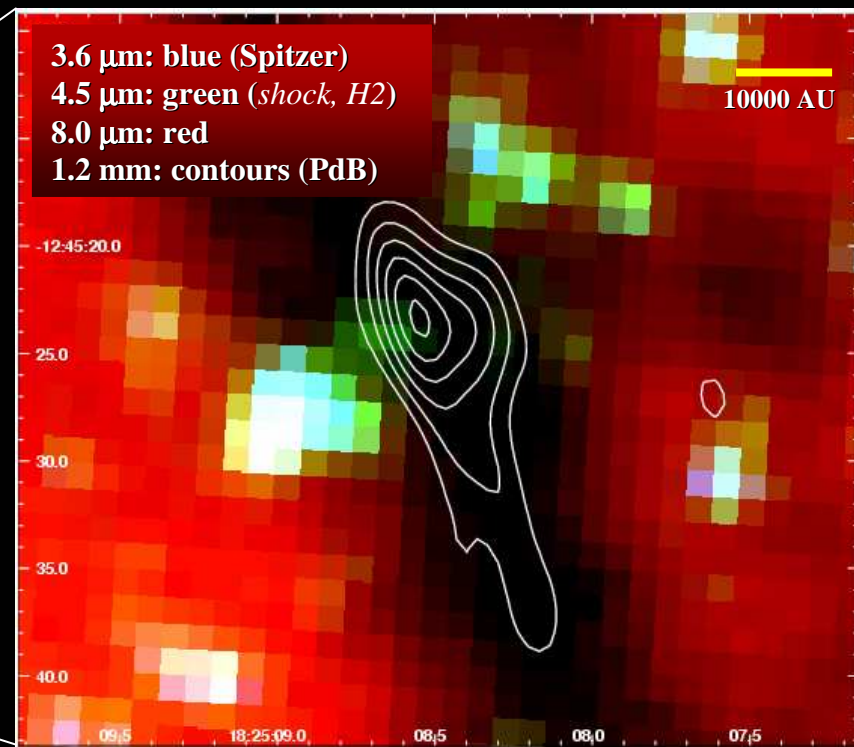
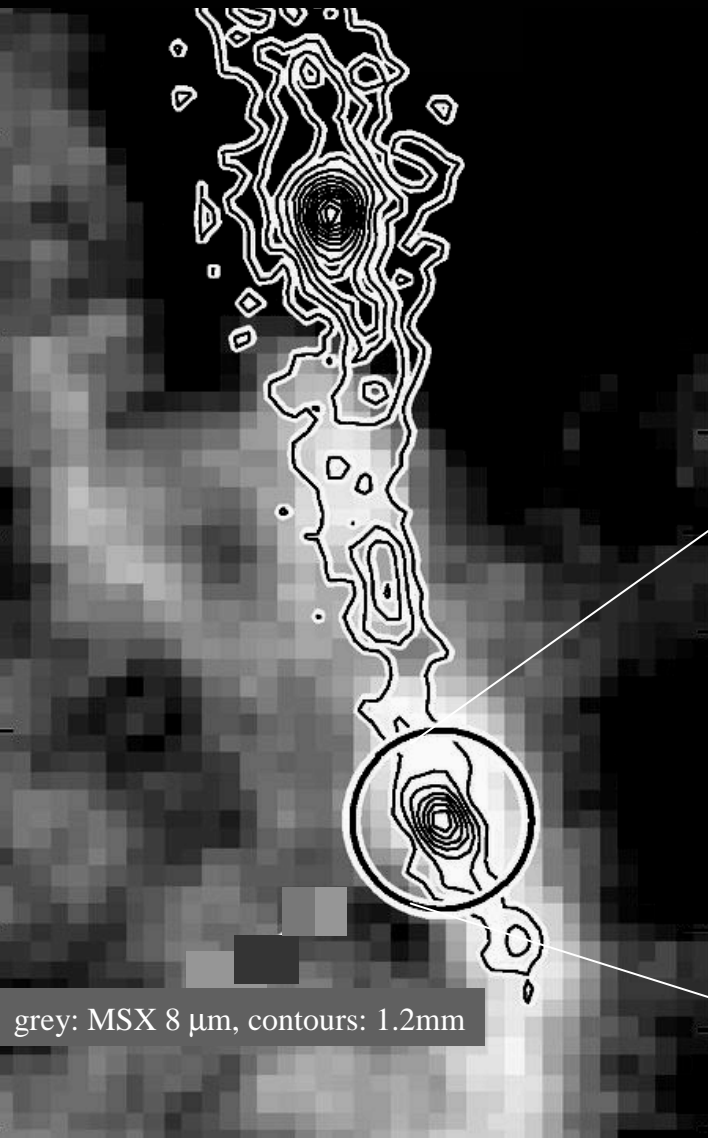
M

$t = 0$:
first hydrostatic core forms

Final stellar mass [M_{sun}]



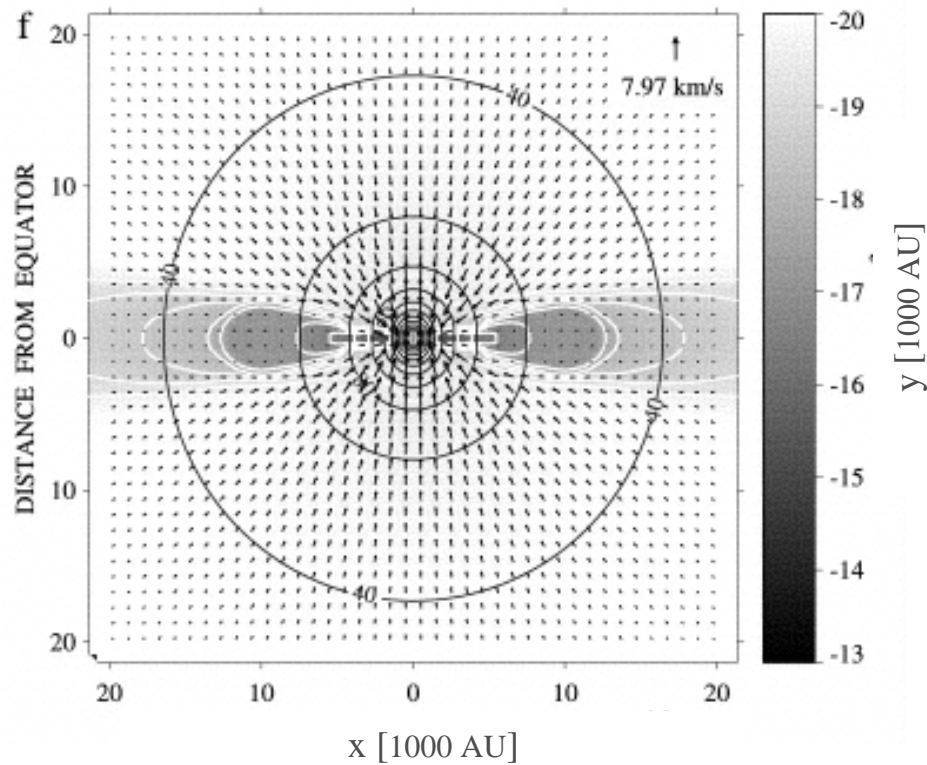
IRDC 18223-3



Disks in HD simulations

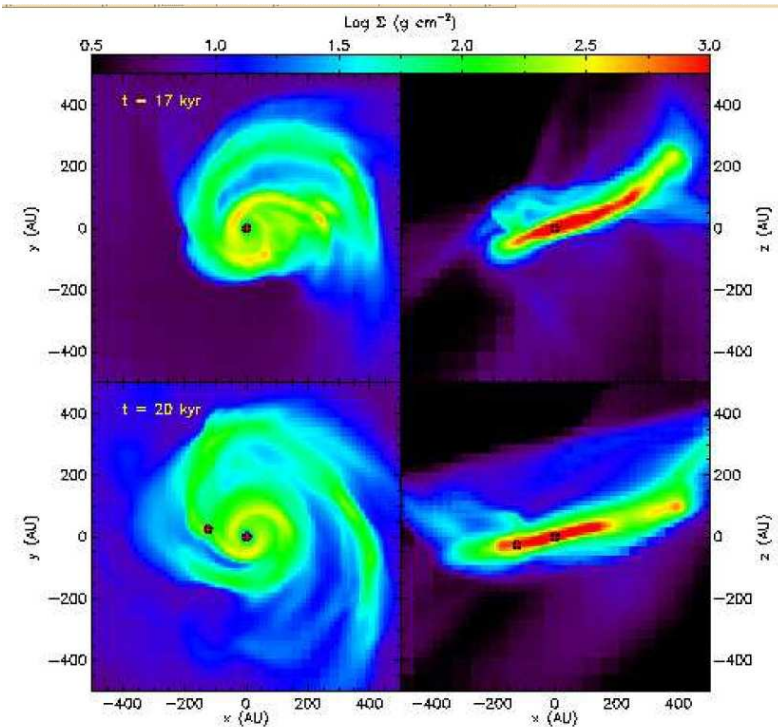
Yorke & Sonnhalter 2002

2D HD + λ -dependent RT



Krumholz, Klein & McKee 2006
ORION MHD AMR

3D HD + flux-limited diffusion



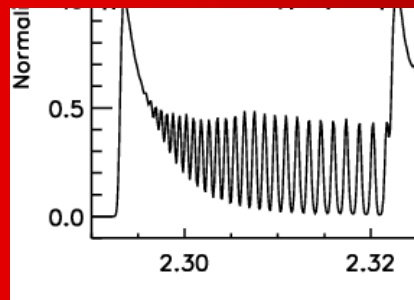
Observing a massive disk candidate

	Opt.	NIR	MIR	FIR	mm	λ
Spatial Resolution:	0.1 kAU		kAU		few kAU	
Continuum:		$n(r,z)$ scattering	$n(r)$ $T(r)$		total disk mass M_d disk extend r_o $n(r,z)$ $T(r,z)$	
Lines:	accretion sign. stellar mass, jets morphology		inner velocity field disk/stellar mass $n(r,z)$ $T(r,z)$ $v(r,z)$		outer velocity field disk/stellar mass, jets $n(r,z)$ $T(r,z)$ $v(r,z)$	

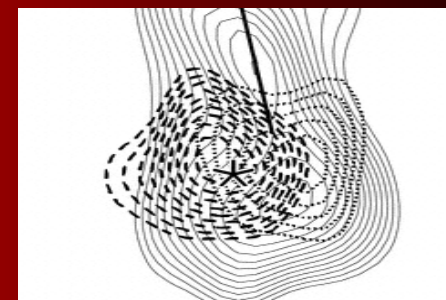
SO-1 K-band
Chini et al. 2004
Steinacker et al. 2006
> 5 Msun 2.4 kpc



IRAS 11097-6102 CO bandhead
Bik & Thi 2004
11 Msun 2.8 kpc



HCOOCH3 in IRAS 18089-1732
Beuther, Zhang, Sridharan, Chen 2005
16 Msun 3.6 kpc



Ray-tracing at high optical depth

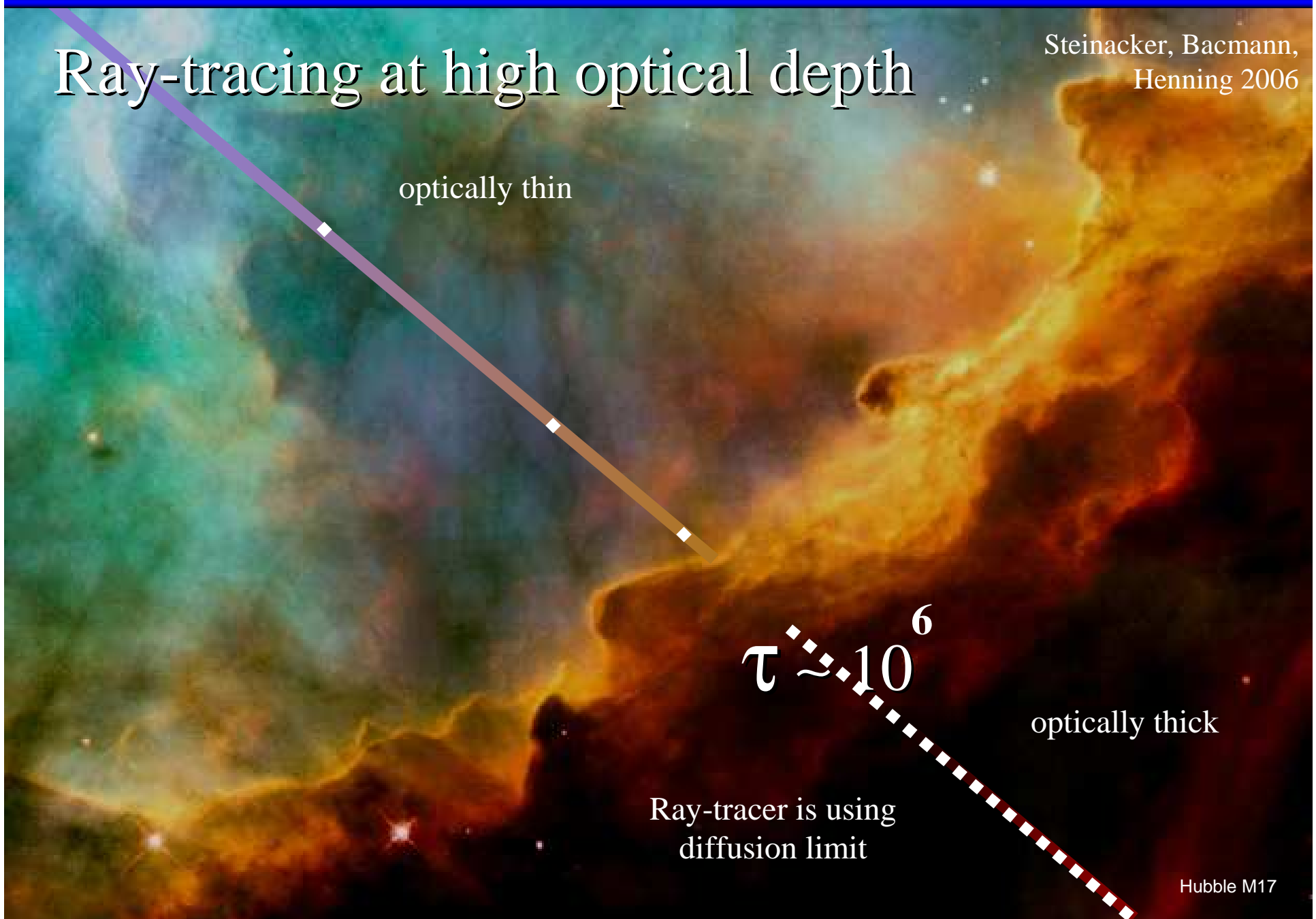
Steinacker, Bacmann,
Henning 2006

optically thin

$$\tau \sim 10^6$$

optically thick

Ray-tracer is using
diffusion limit



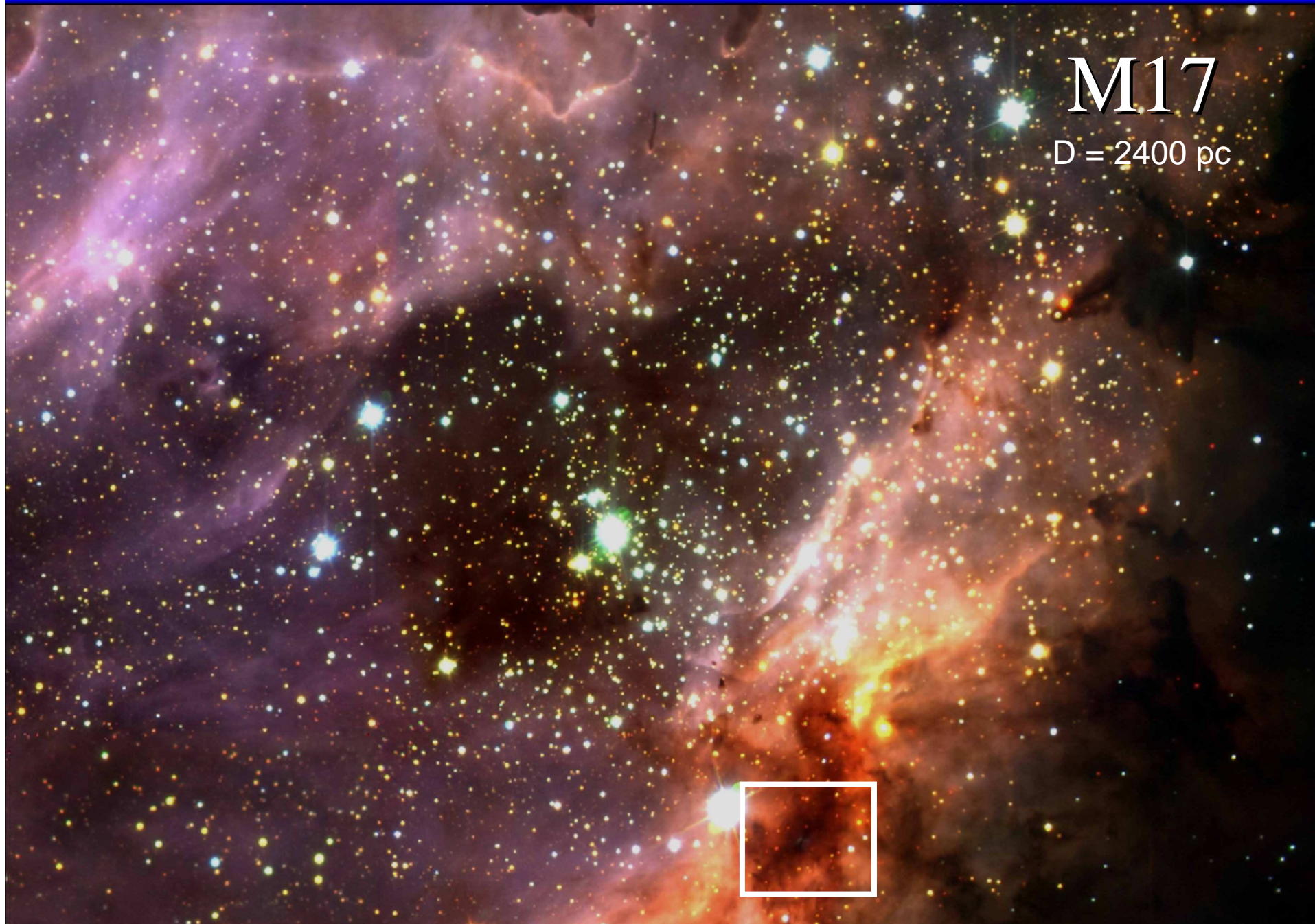
M17

D = 2400 pc



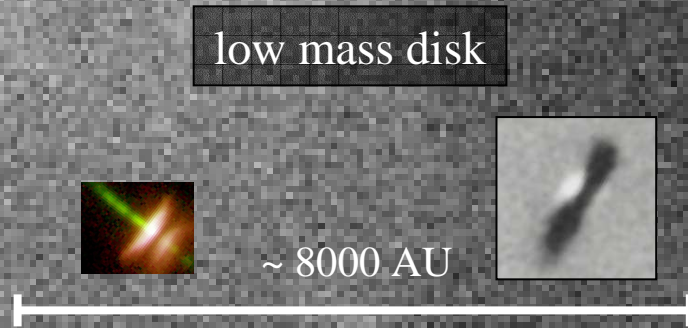
M17

D = 2400 pc



SO-1





SO-1

Chini et al. Nature 2004

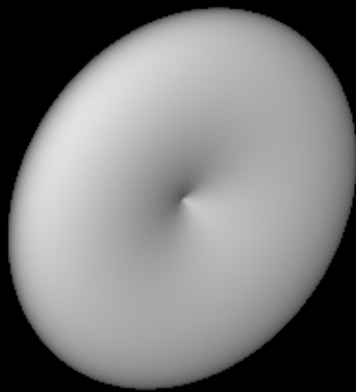
ISAAC/VLT 2.2 μm

SO-1

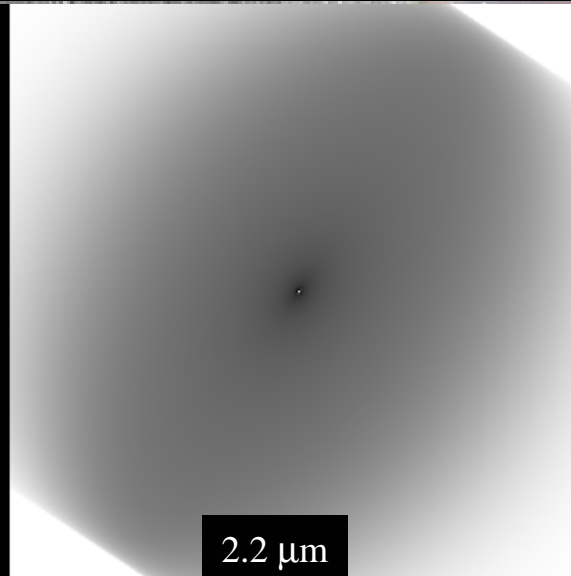
Steinacker, Chini, Nielböck, Nürnberger,
Hoffmeister, Hure, Semenov 2006

Disk mass:
0.2-13 solar masses

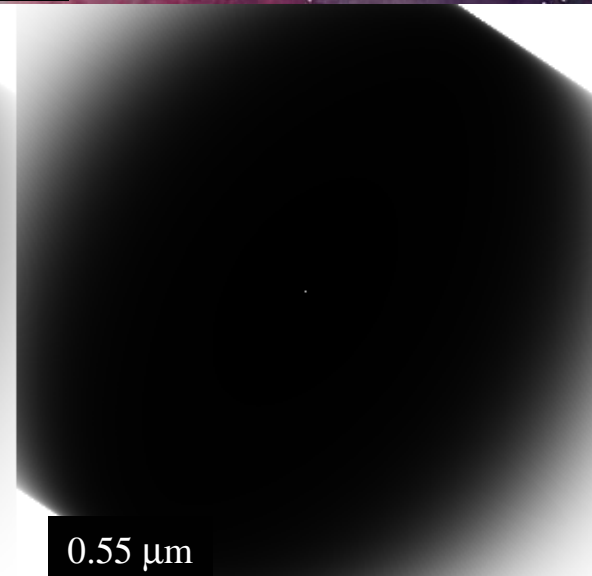
8 parameter circumstellar disk fit



Iso-density surface



2.2 μm



0.55 μm

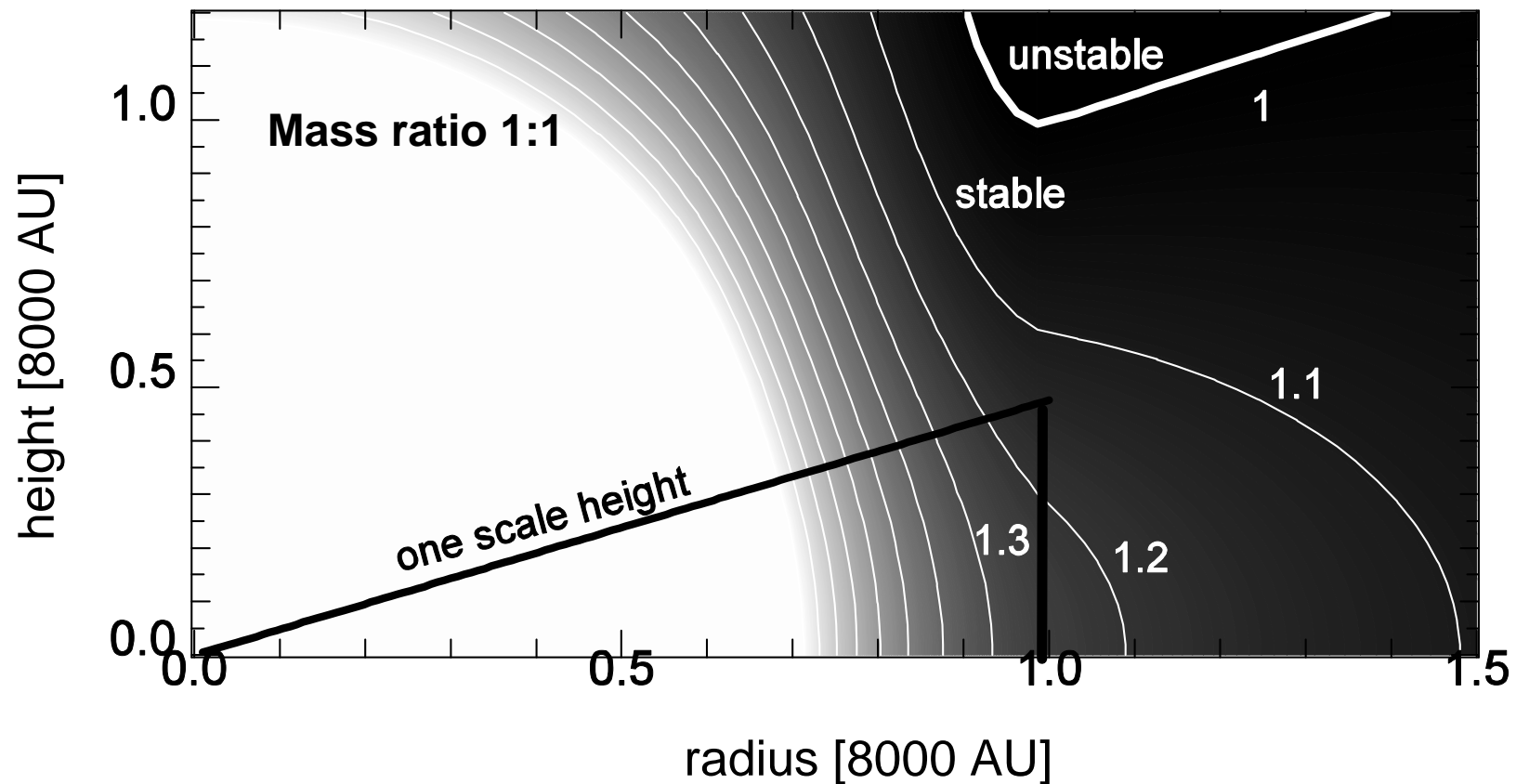
Steinacker, Chini, Nielbock, Nürnberger, Hoffmeister, Hure, Semenov 2006

Given both star and disk are massive:

Is such a disk stable?

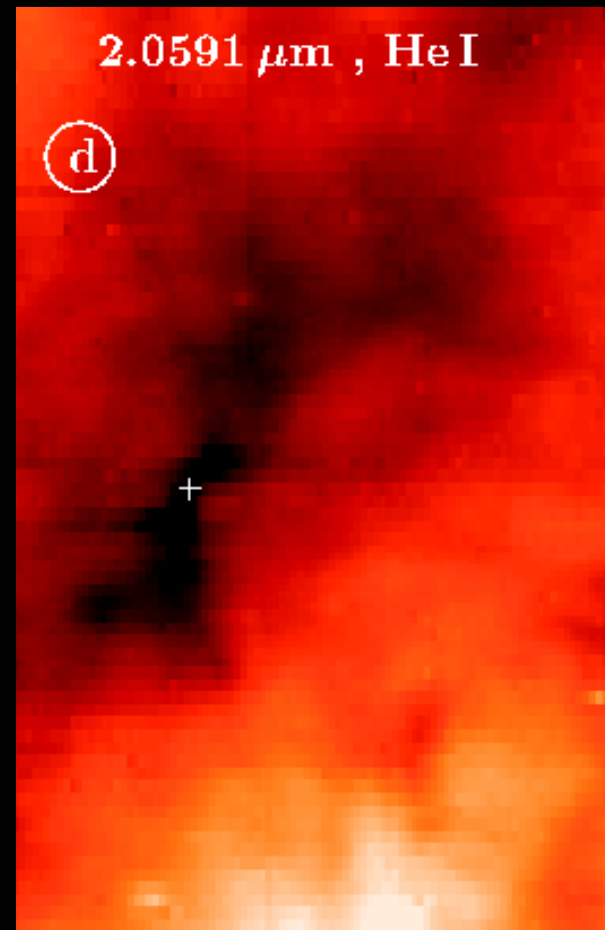
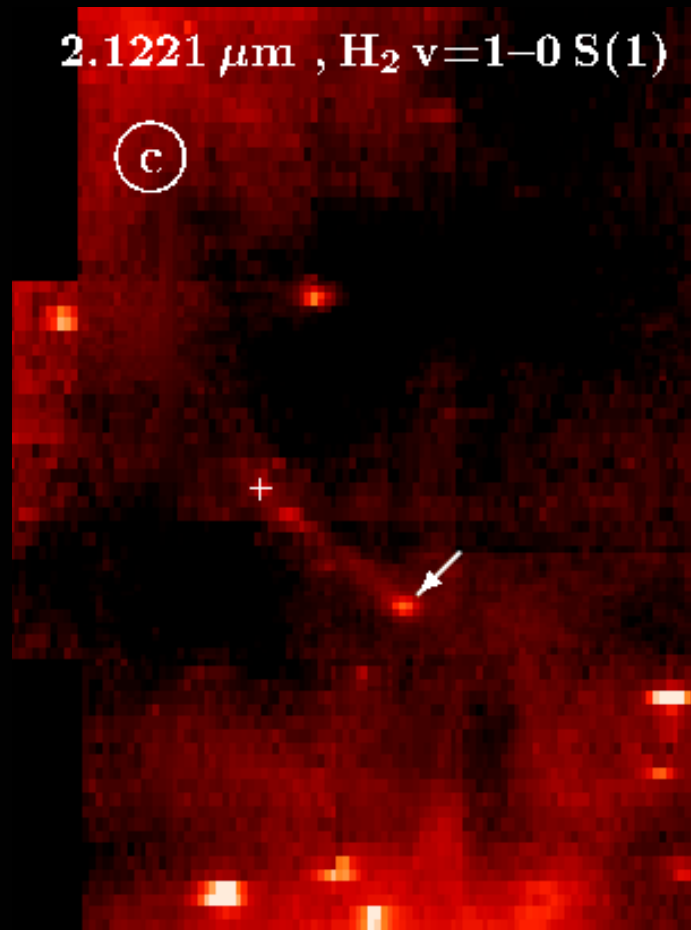
 $\frac{\text{central gravitational acceleration}}{\text{disk gravitational acceleration}}$

disk gravitational acceleration



Collimated jet confirmed

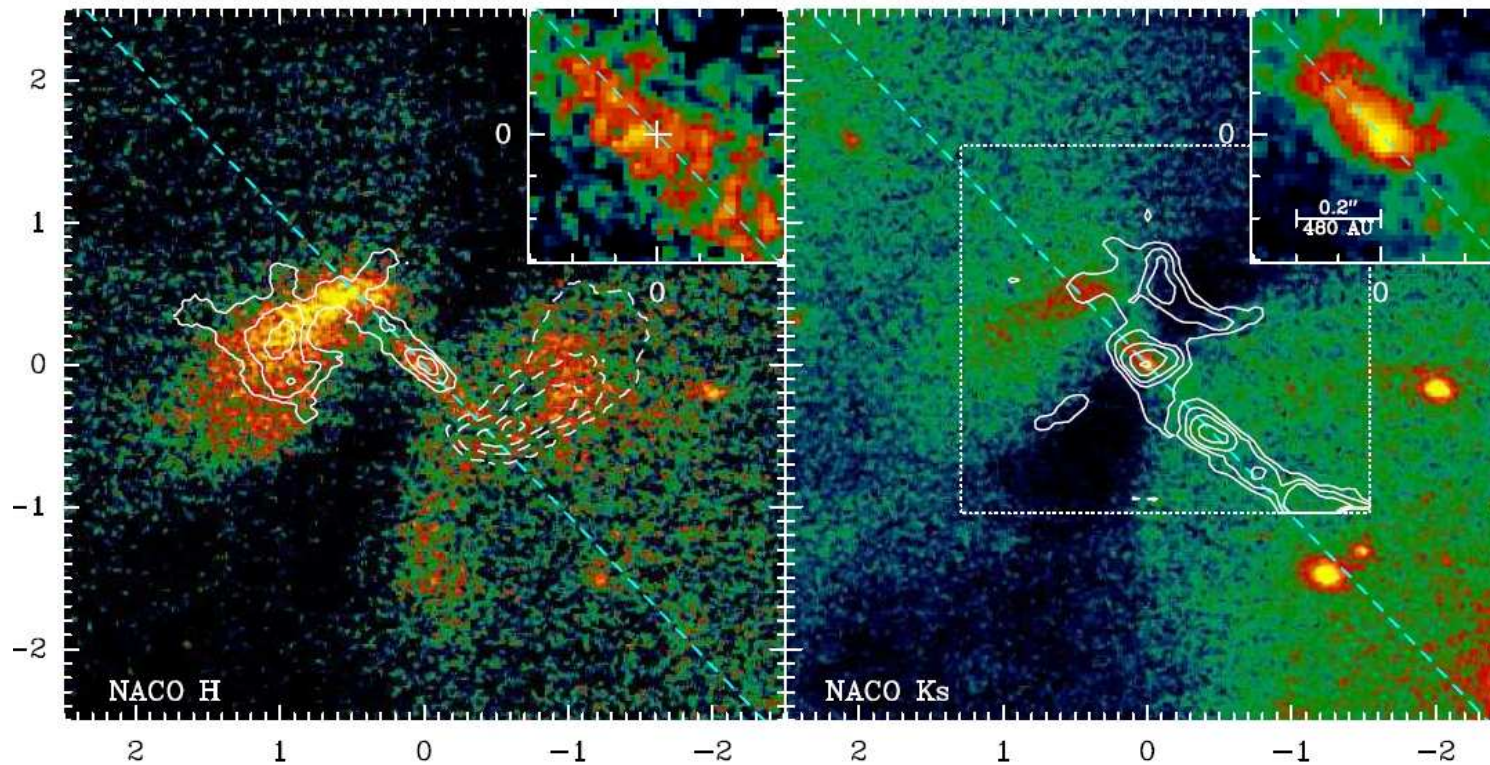
Nürnberger et al. 2007



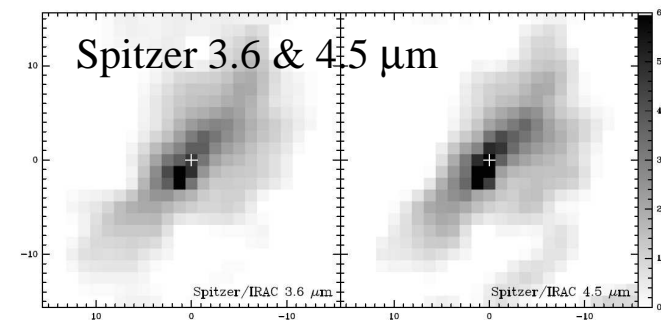
K band integral field spectroscopy SINFONI, VLT

Nielbock, Chini, Hoffmeister, Nürnberger, Scheyda, Steinacker 2007

Central region resolved down to 240 AU

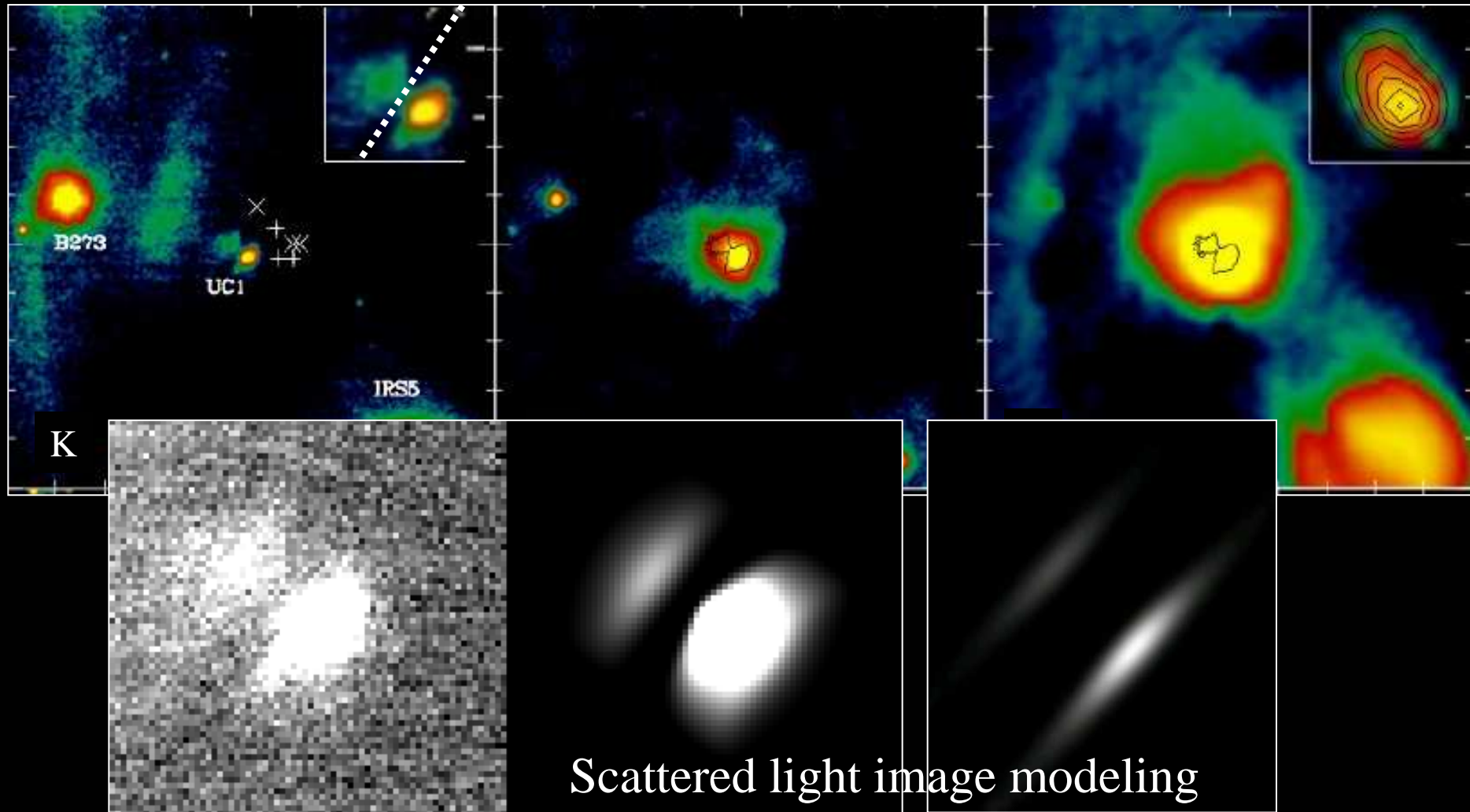


Extinction map from Spitzer as well as H₂ emission at the disk surface point towards > B4 star



Nielbock, Chini, Hoffmeister, Scheyda, Steinacker, Nürnberger, Siebenmorgen 2006

UC1 – A *hyper*-compact HII region



Chini, Hoffmeister, Nielbock, Scheyda, Steinacker, Siebenmorgen, Nürnberger 2006

A remnant disk around a massive star (B0)

ISAAC NTT 3.8 μm

TIMMI2 ESO3.6 10 μm

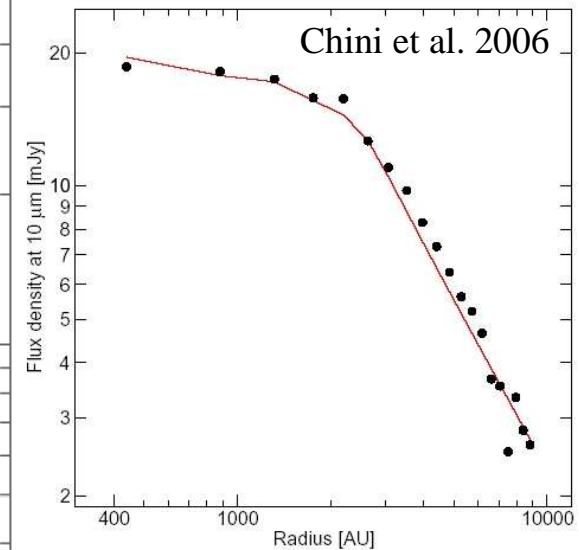
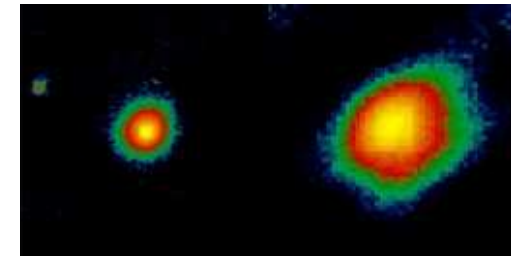
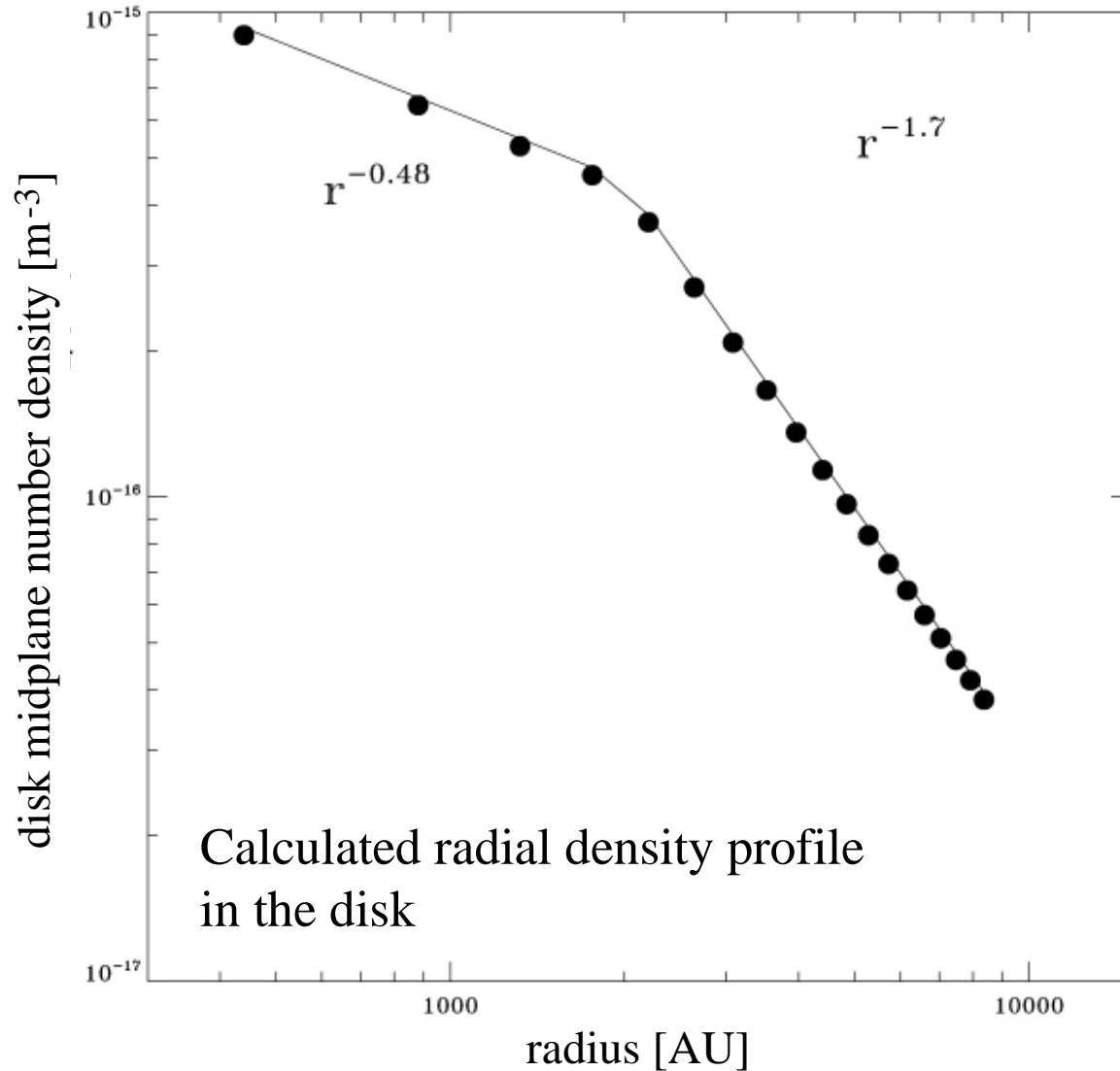
26 solar masses

Pa α and He I 10314 line: B0V to B1V

10000 AU

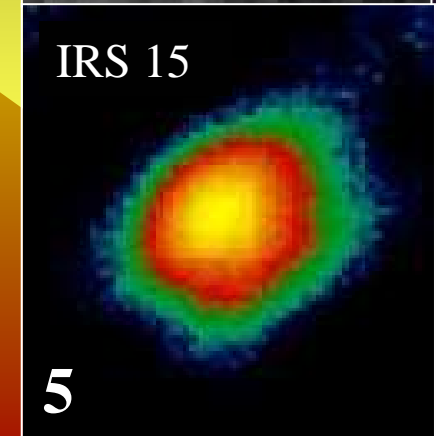
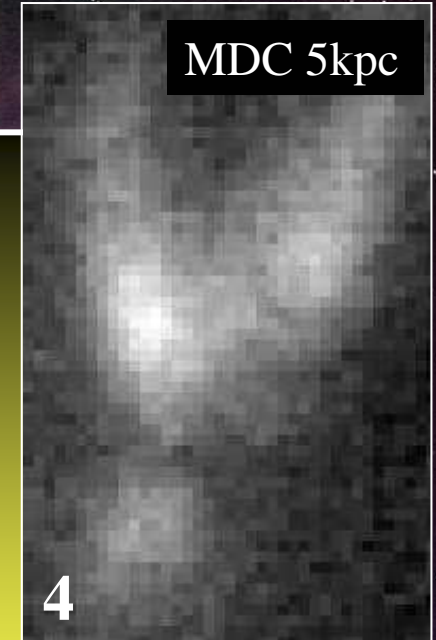
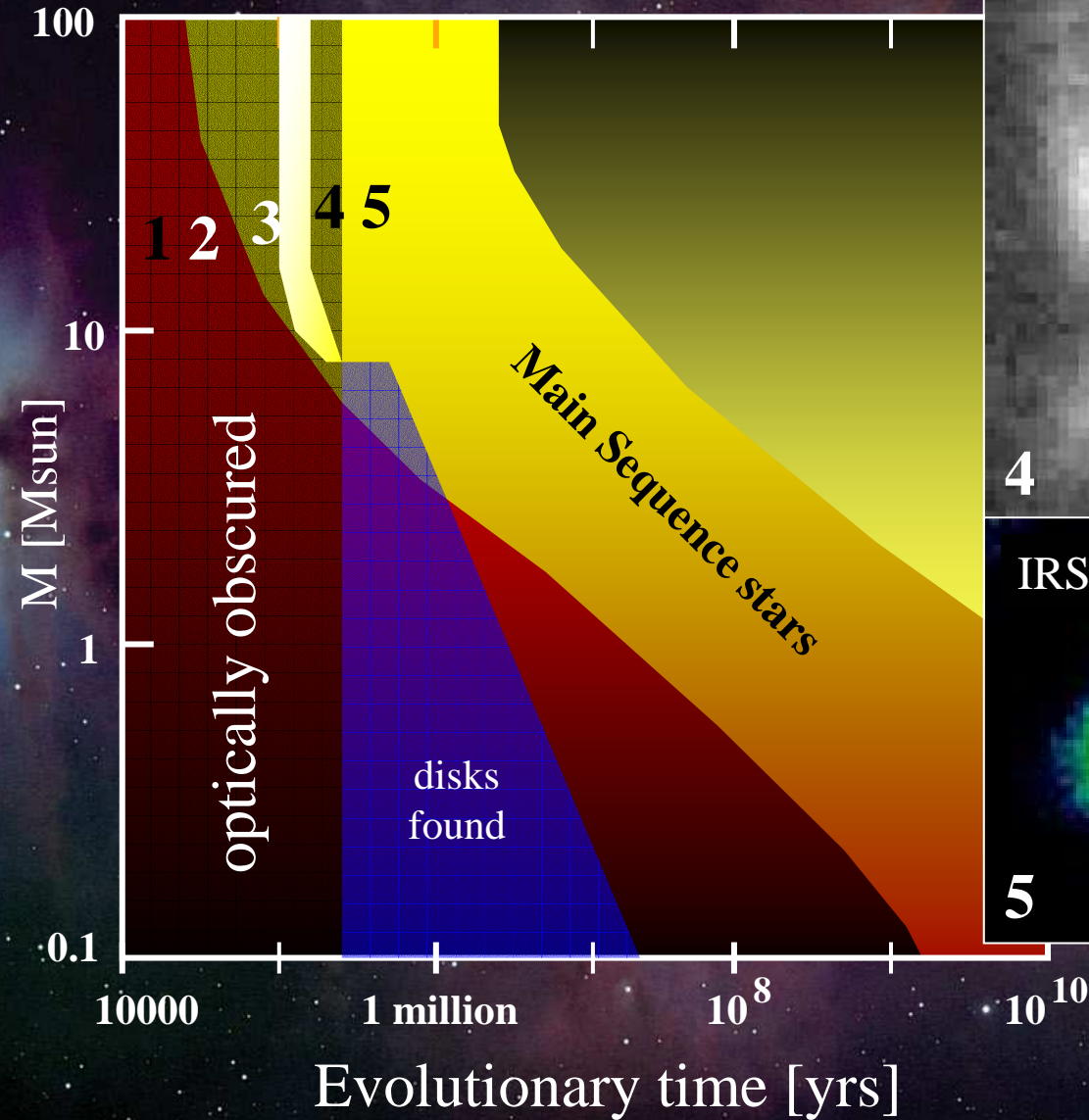
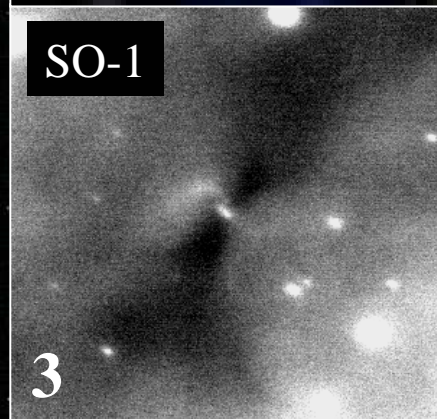
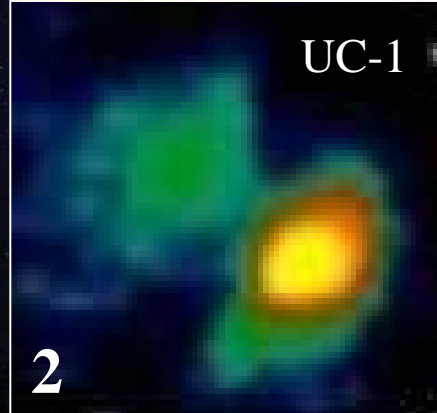
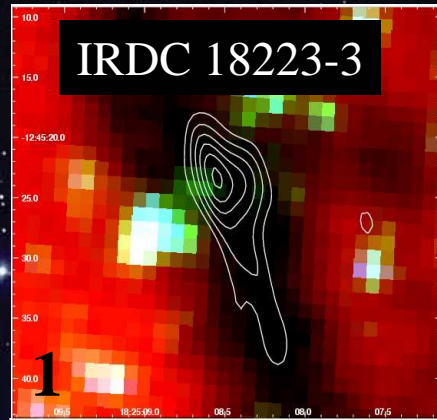


Indication for an inner cavity



Growing evidence for disk-like structures around Young Massive Stars

Jürgen Steinacker



Cosmic Dust Near & Far

8.-12. September 2008
Convention Center, Heidelberg, Germany

Scientific Organizing Committee:

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www.mpia.de/DNF08



8-12 September
2008

Convention Center
Heidelberg

Follow-up workshop to Cosmic Dust – Near & Far:



Cosmic Dust & Radiative Transfer

a workshop devoted to radiative transfer coding

15-17 Sep 2008
MPIA, Heidelberg, Germany
(by invitation only)

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stein@mpia.de swolf@mpia.de