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CAMBRIDGE

Constraining reionization with Lyman-alpha emitters and the CMB

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KICC / IoA, University of Cambridge

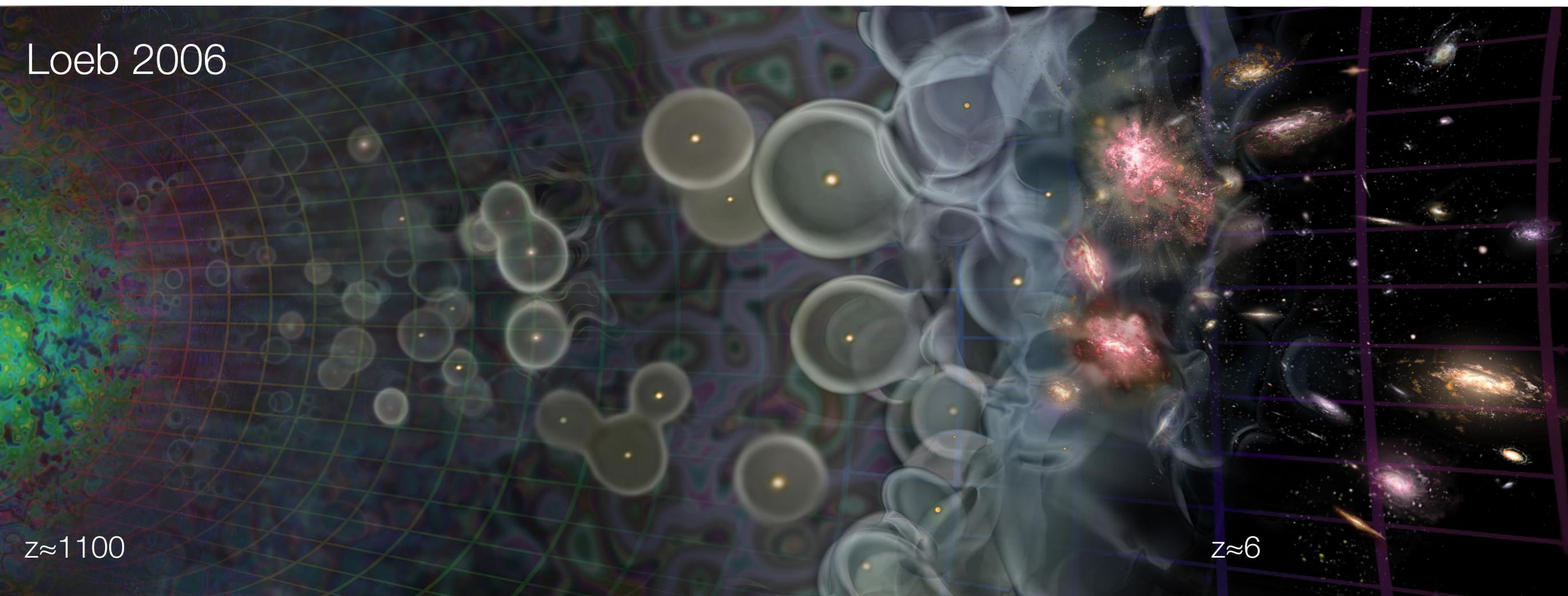
collaborators: Tirthankar Choudhury, Martin Haehnelt, James Bolton

CMB and the epoch of reionization

time



Loeb 2006



$z \approx 1100$

$z \approx 6$

CMB

first stars
and galaxies

today

cosmic
dark ages

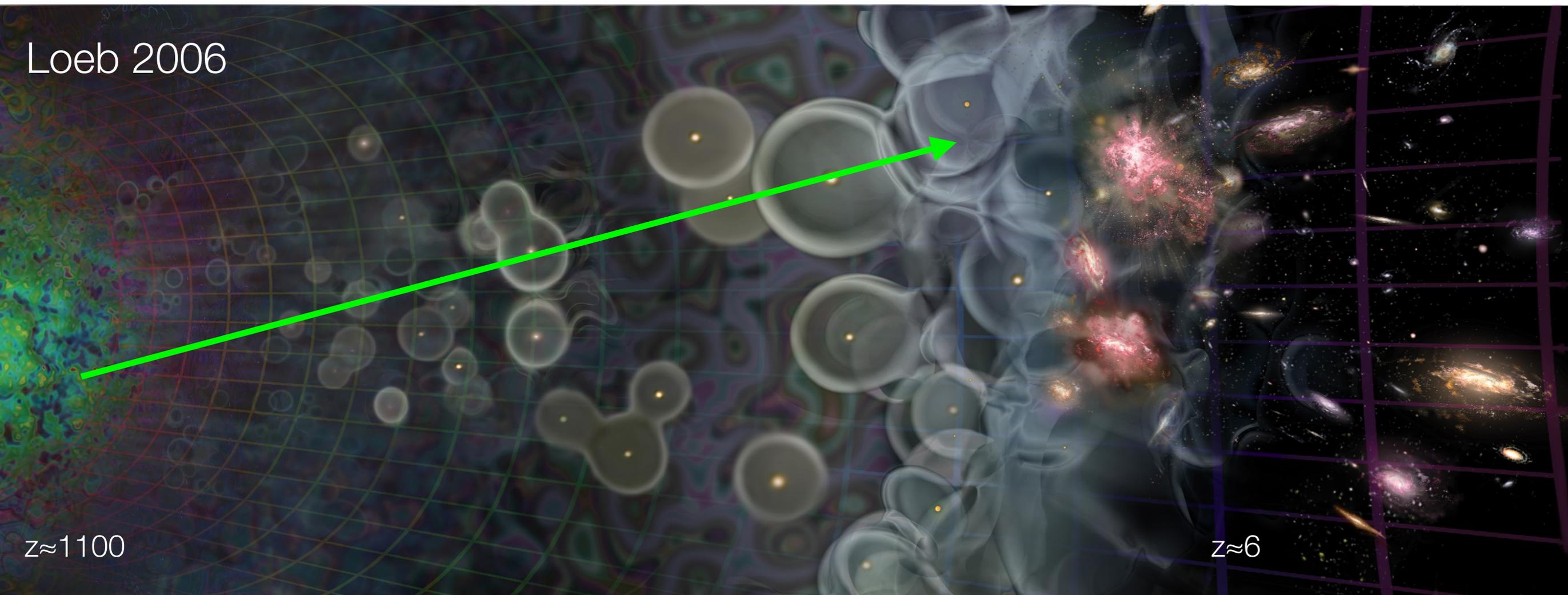
reionization

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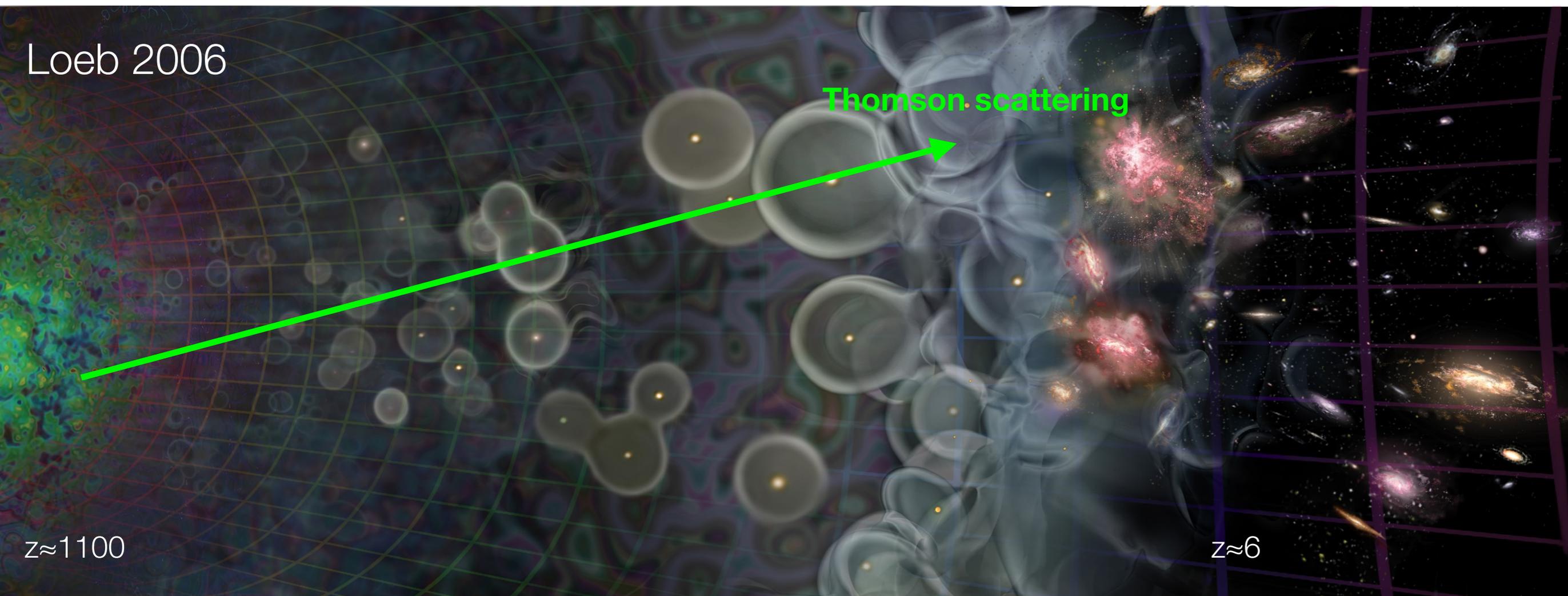
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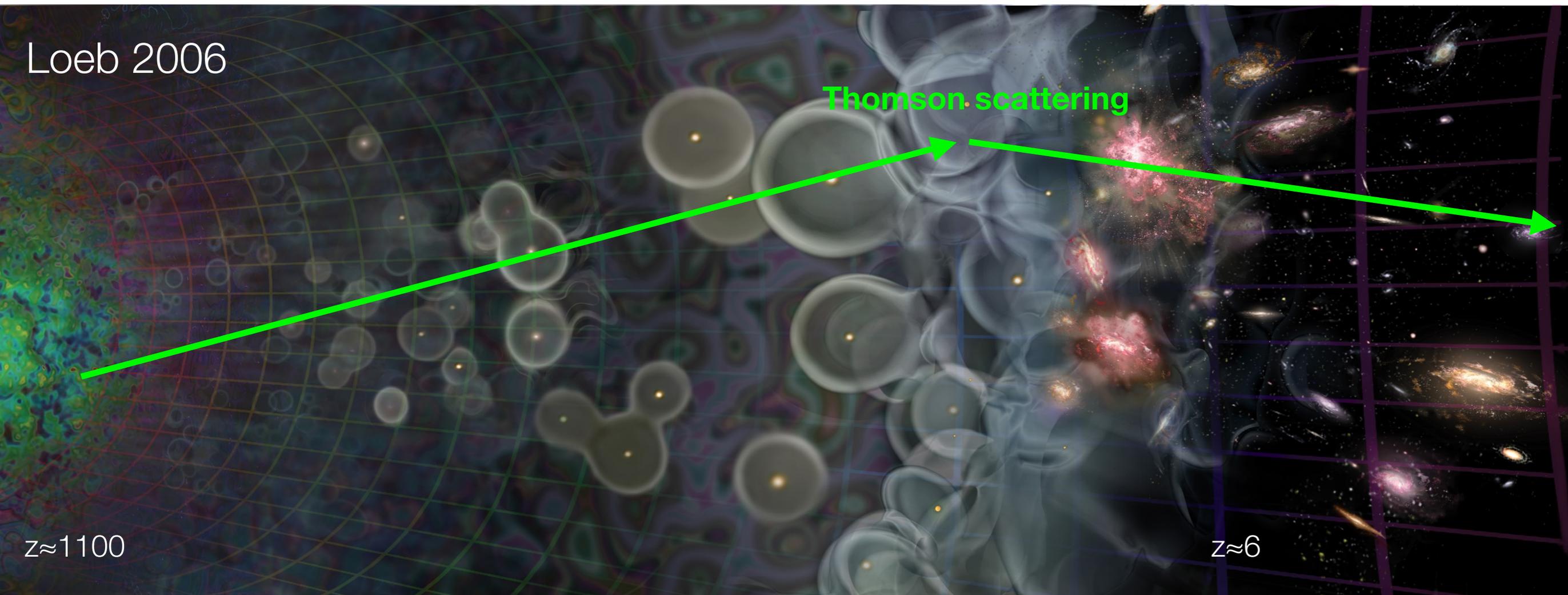
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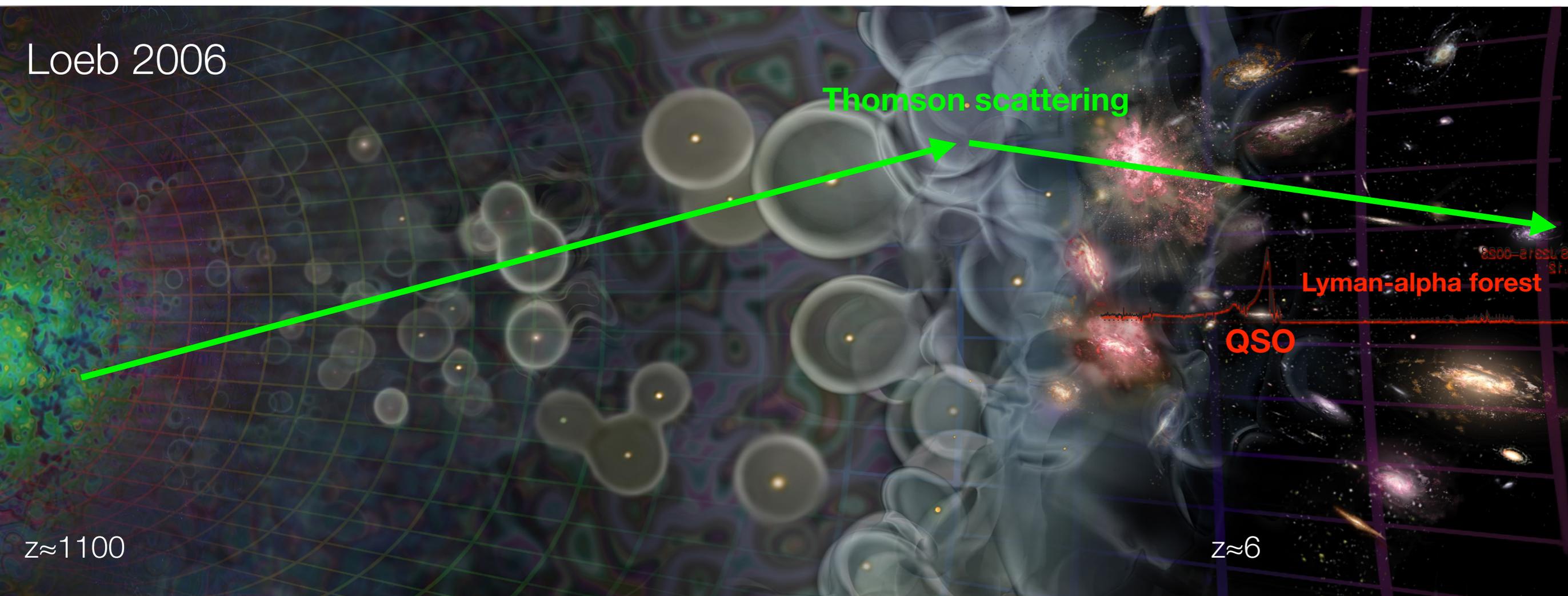
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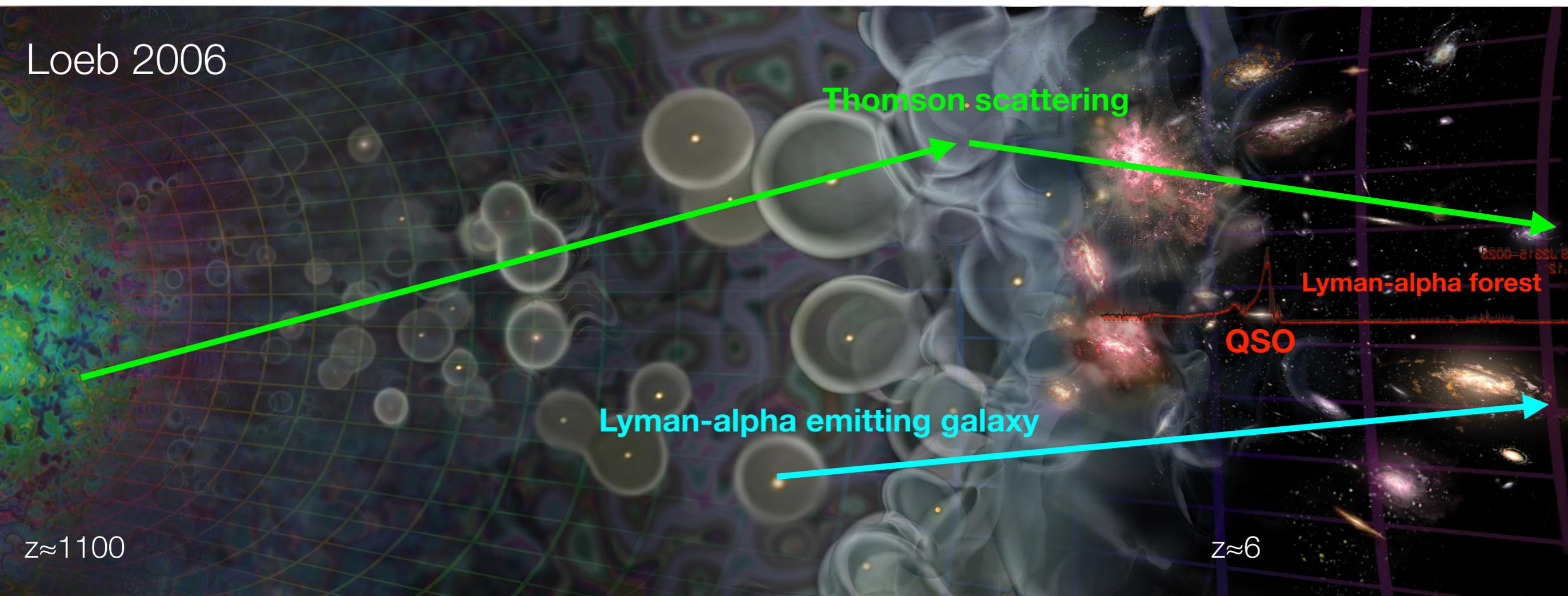
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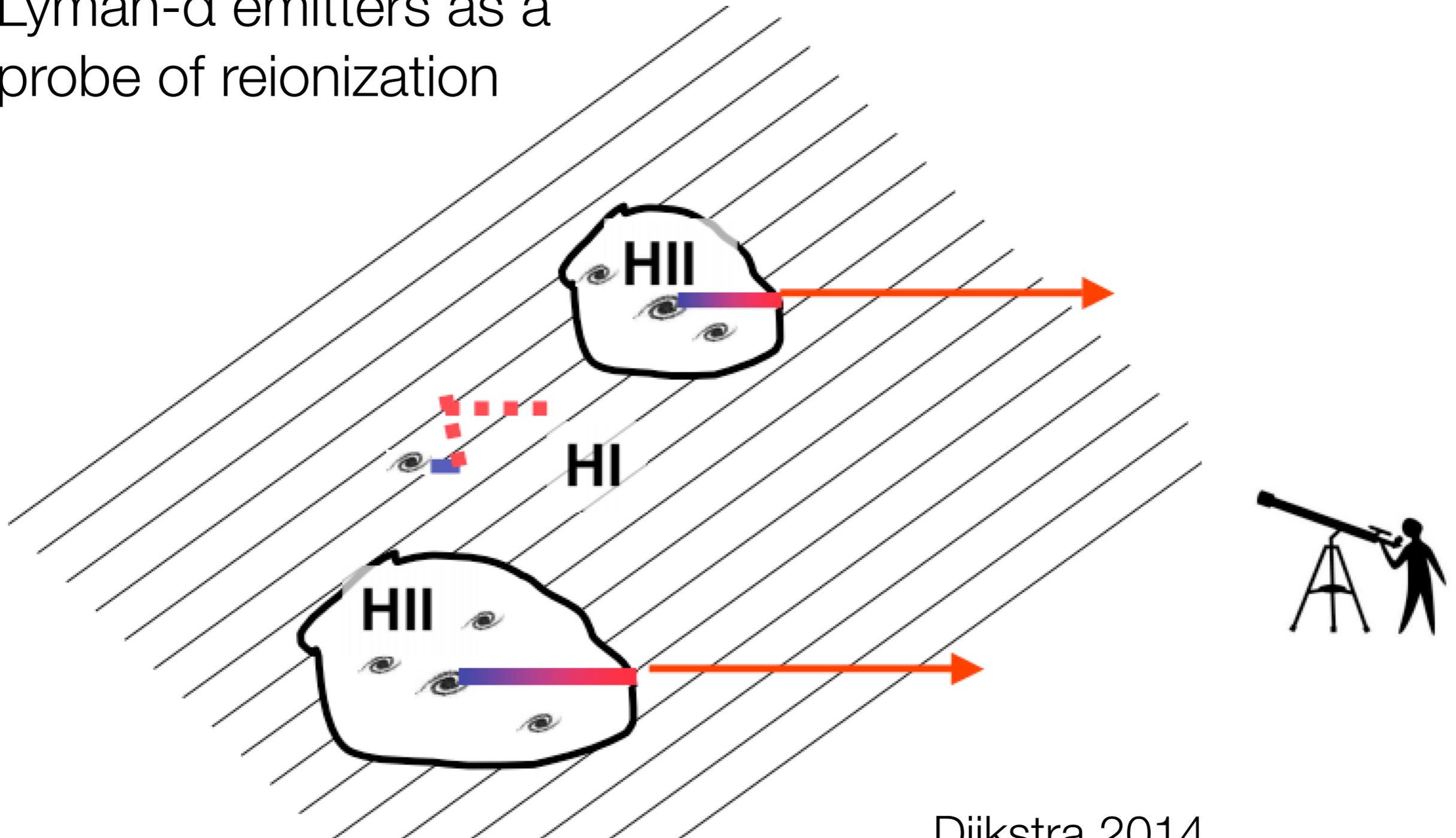
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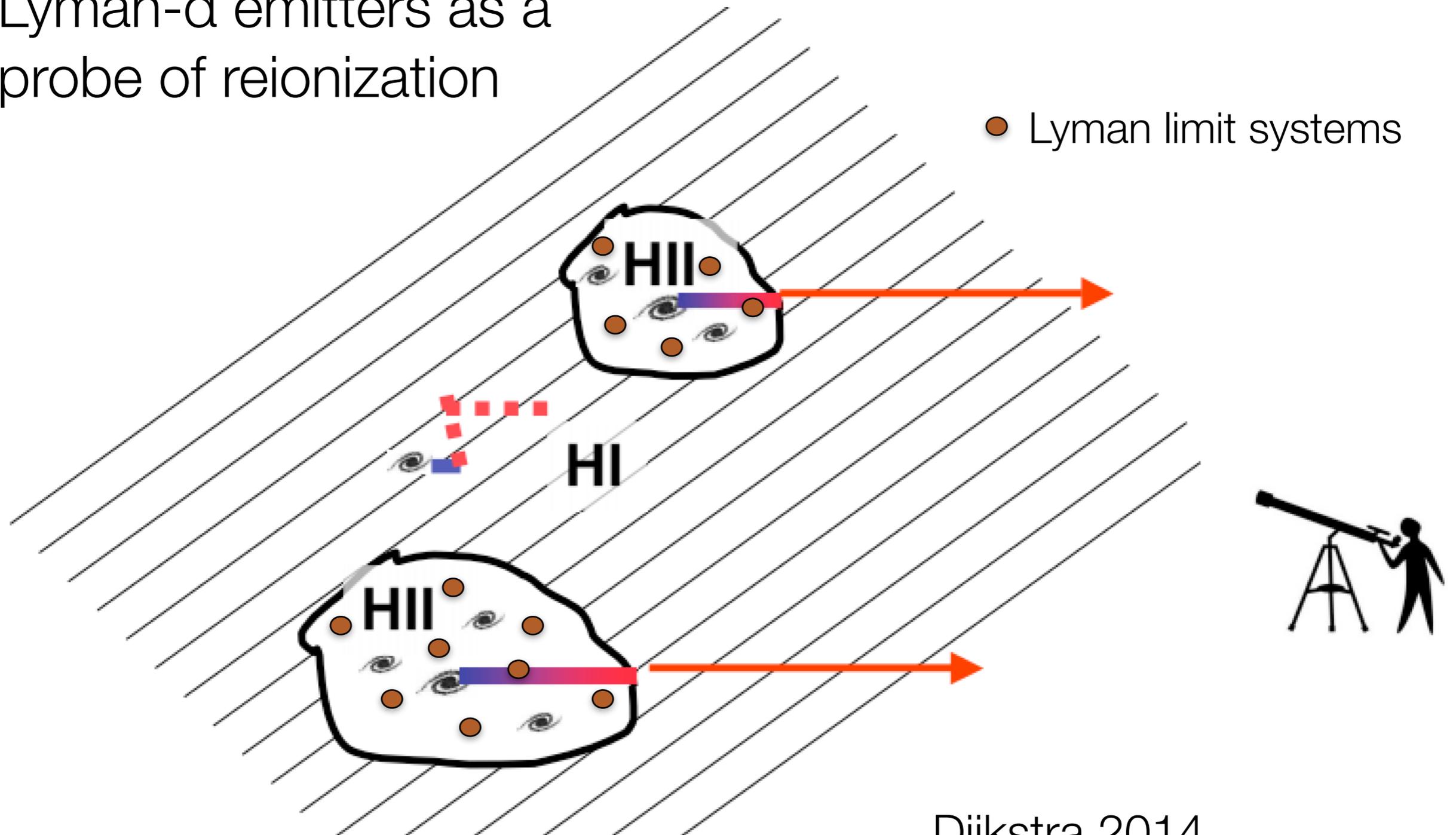
Probing the ionization & thermal state of the IGM

- Lyman- α emitters as a probe of reionization



Probing the ionization & thermal state of the IGM

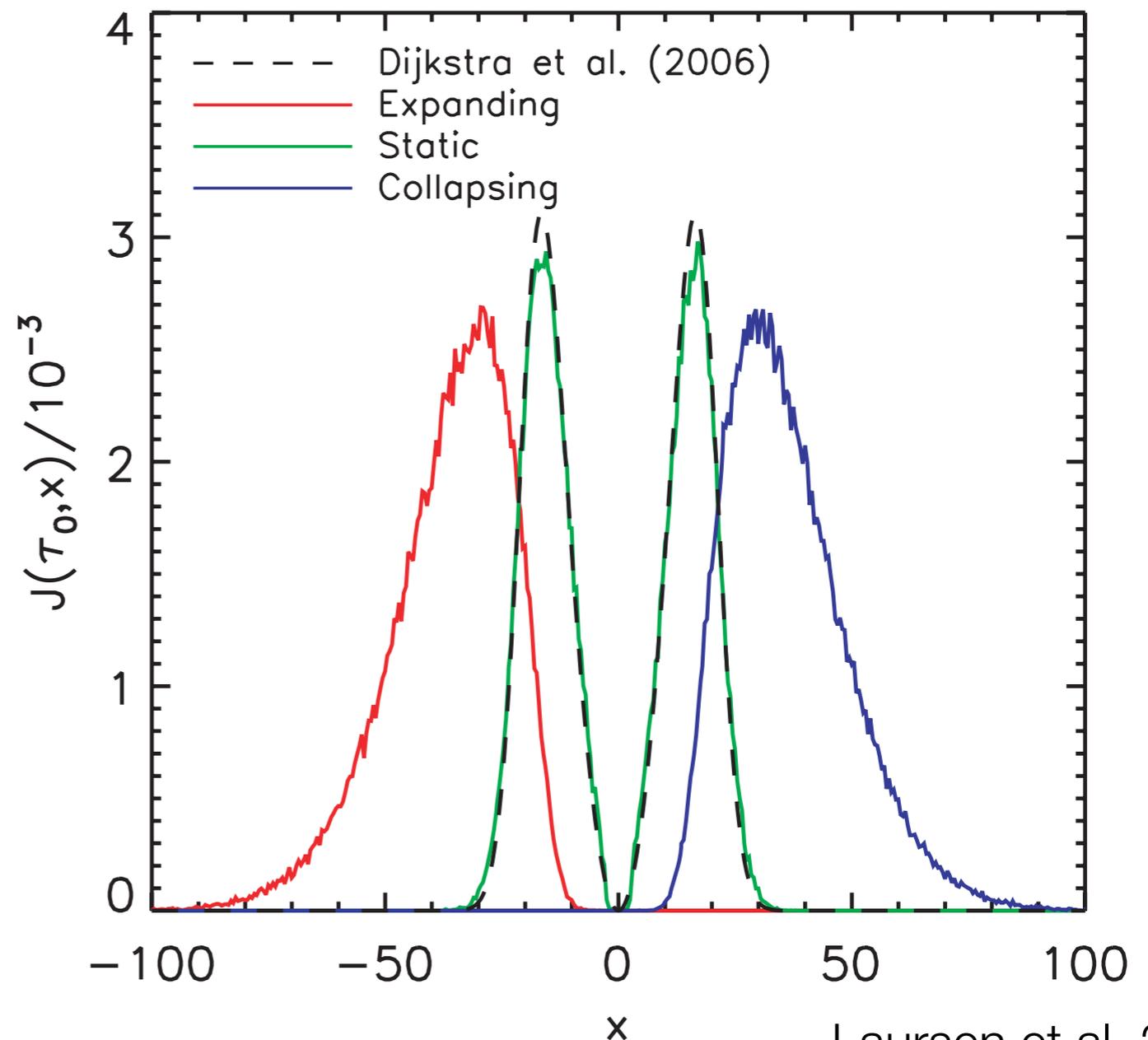
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Probing the ionization & thermal state of the IGM

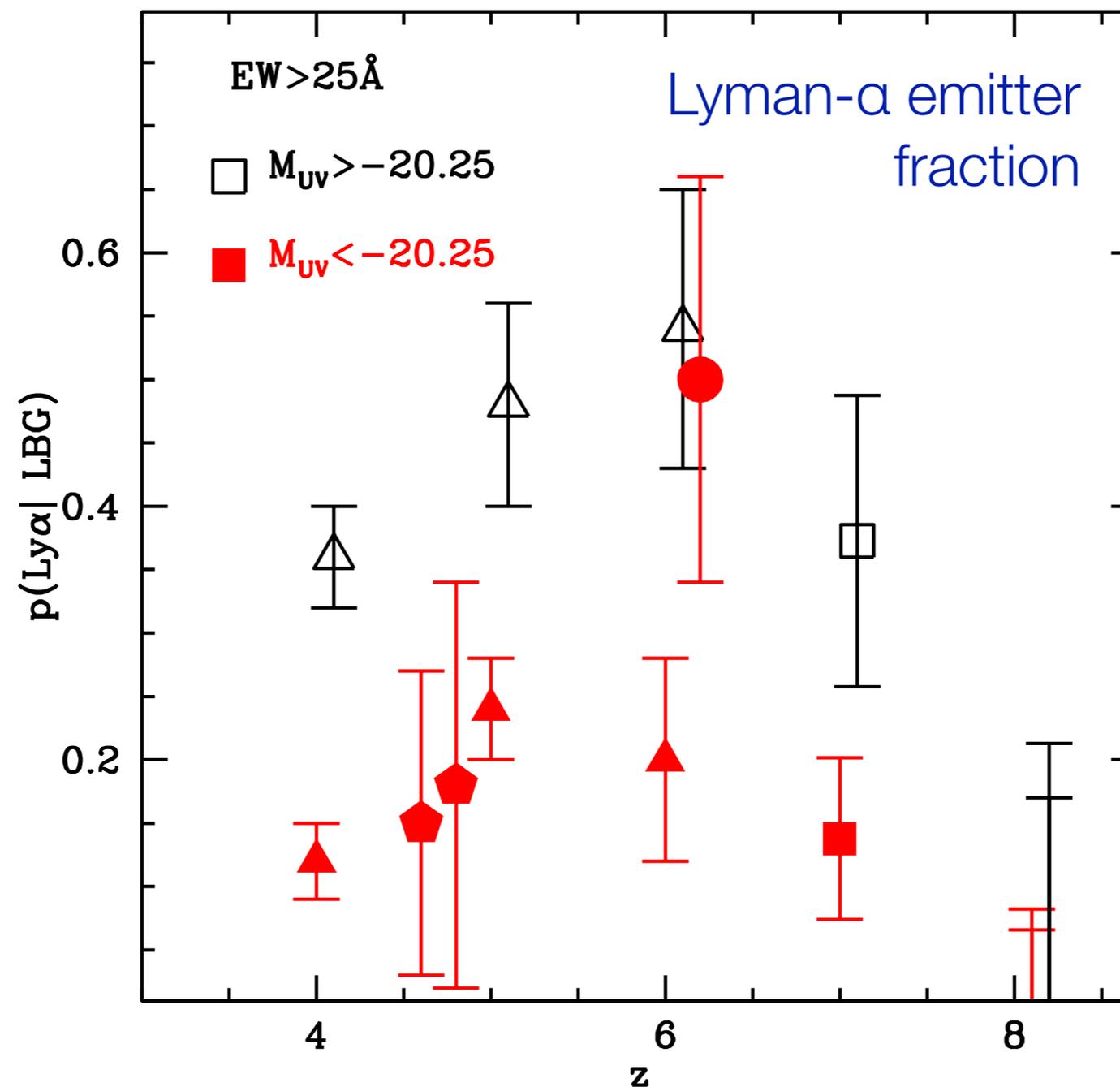
frequency diffusion of Lyman- α photons

- Lyman- α emitters:
- line usually redshifted with respect to systemic velocity



Probing the ionization & thermal state of the IGM

- Lyman- α emitters as a probe of reionization



Semi-numerical models of reionization

- semi-numerical models based on the ionizing photon budget, ionized if:

ionizing photons - recombinations > number of hydrogen atoms

- excursion set approach:

check for each point if there is any radius inside which the ionization condition is satisfied



no



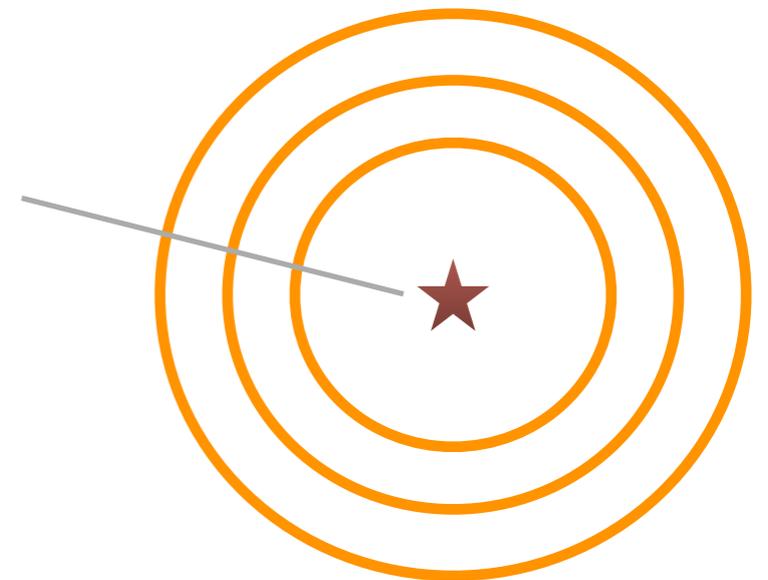
not ionized



yes

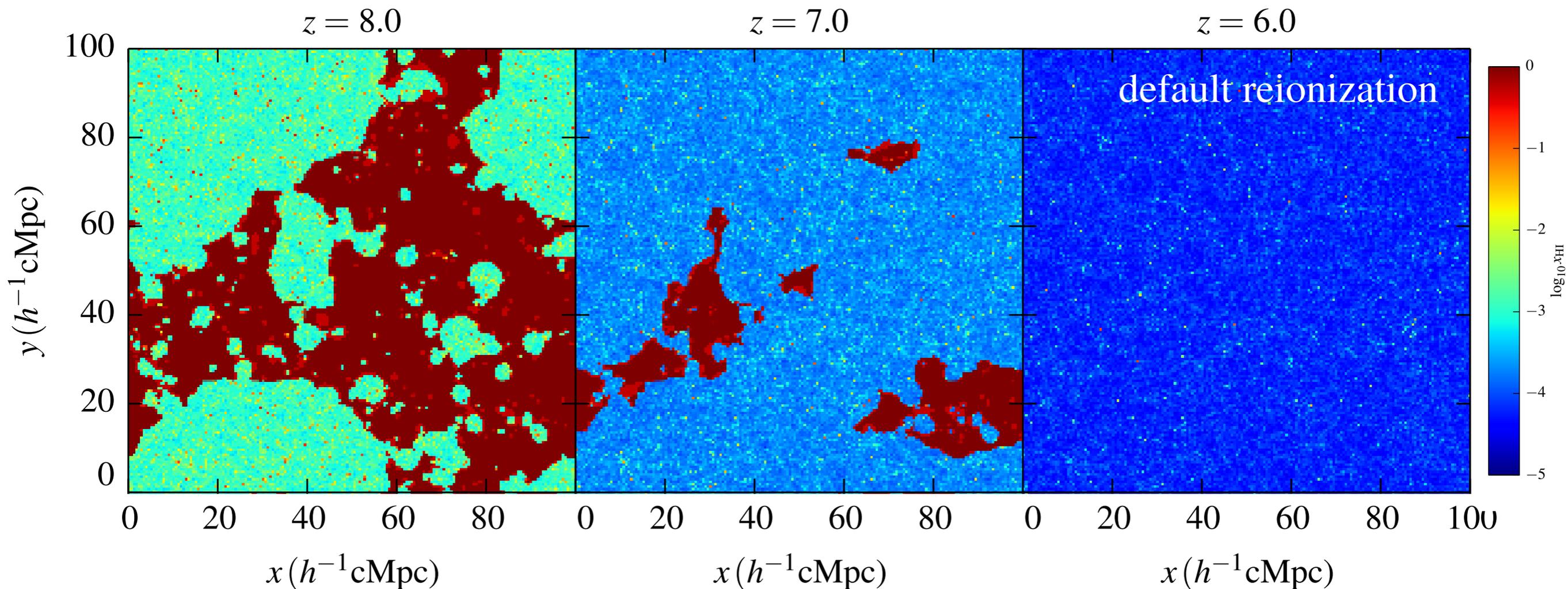


ionized

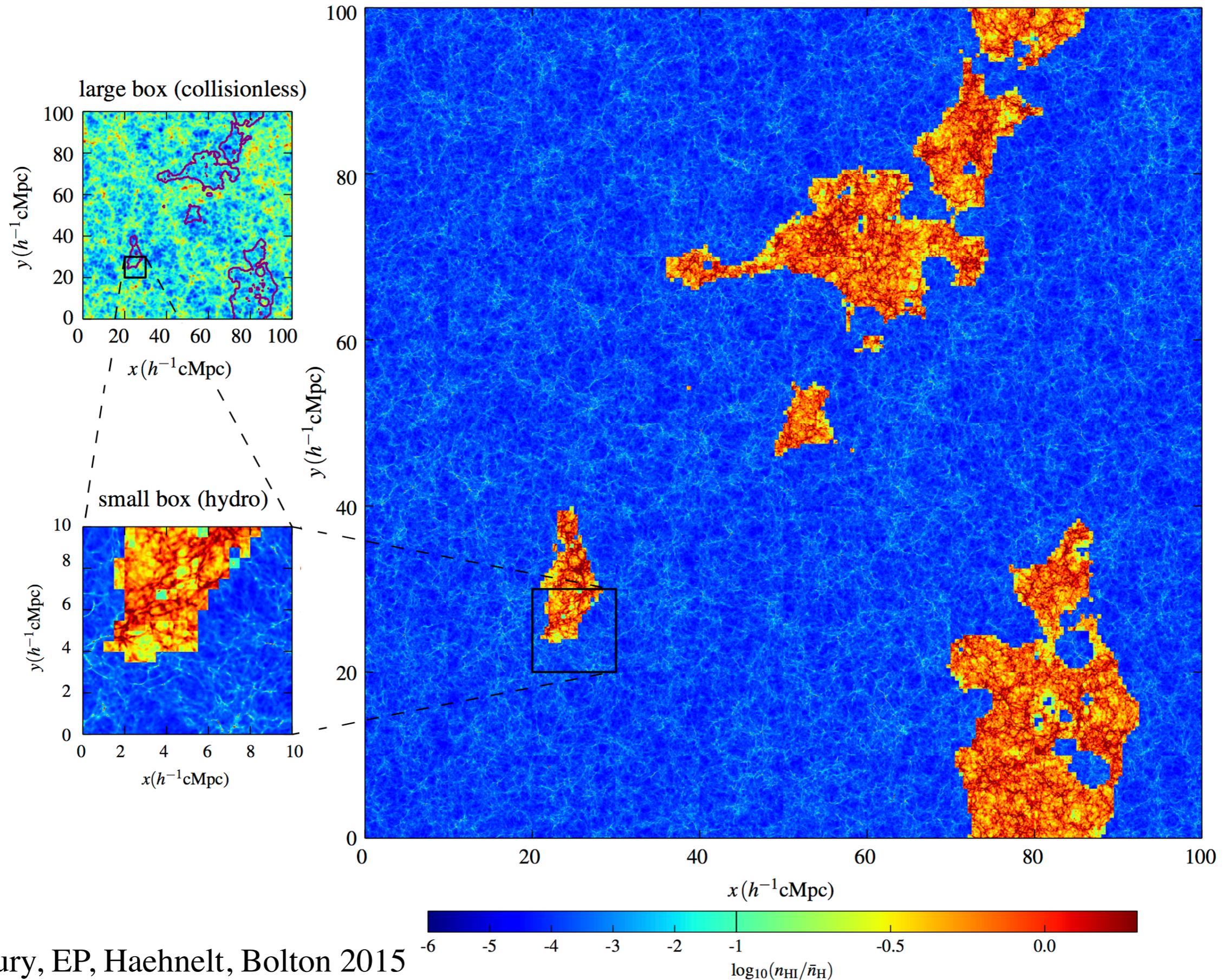


Semi-numerical reionization in post-processing

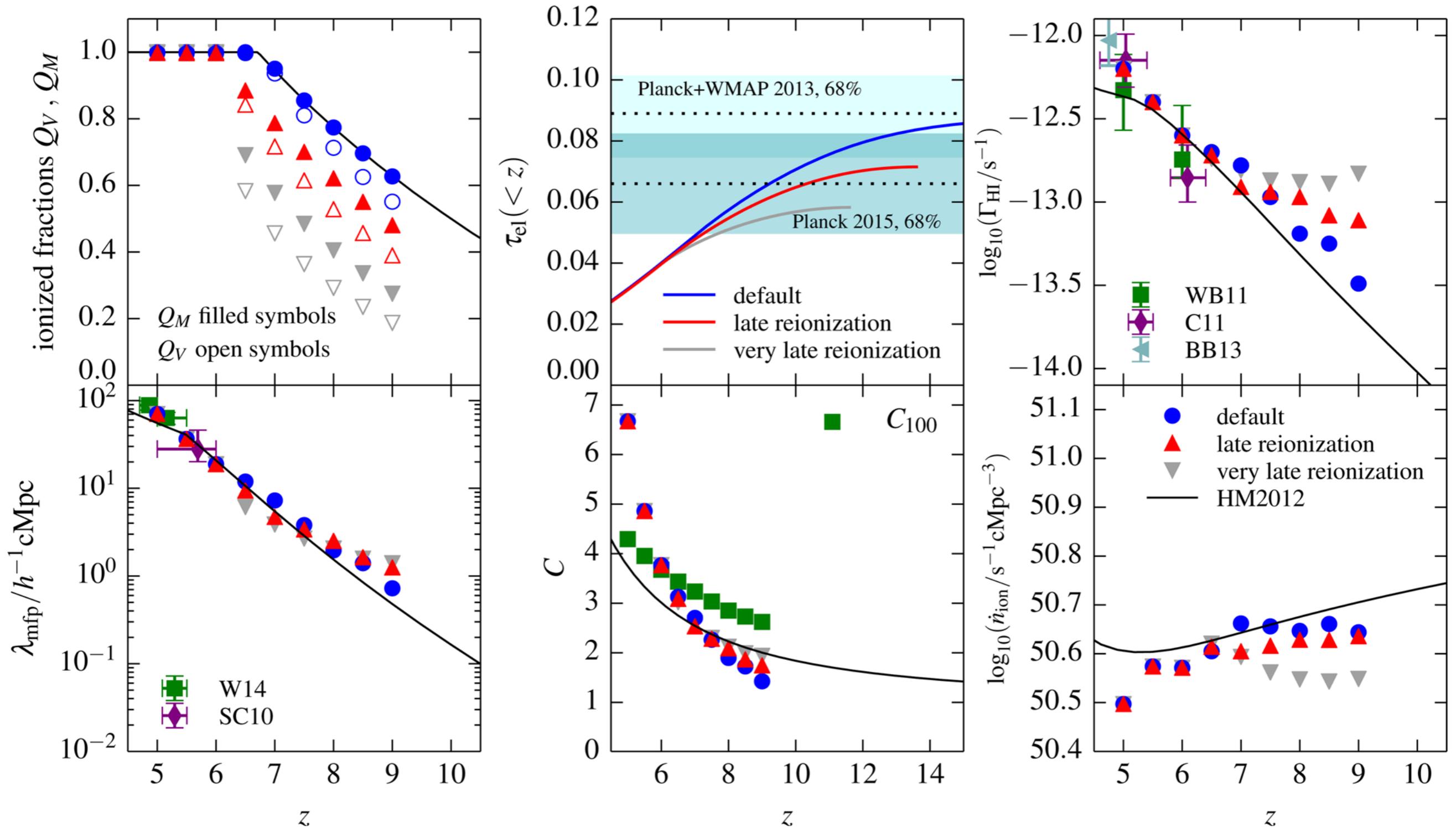
- empirical assignment of ionizing luminosity to halos
- > get ionized regions in post-processing from simulations



hybrid technique combining large and small box hybrid box

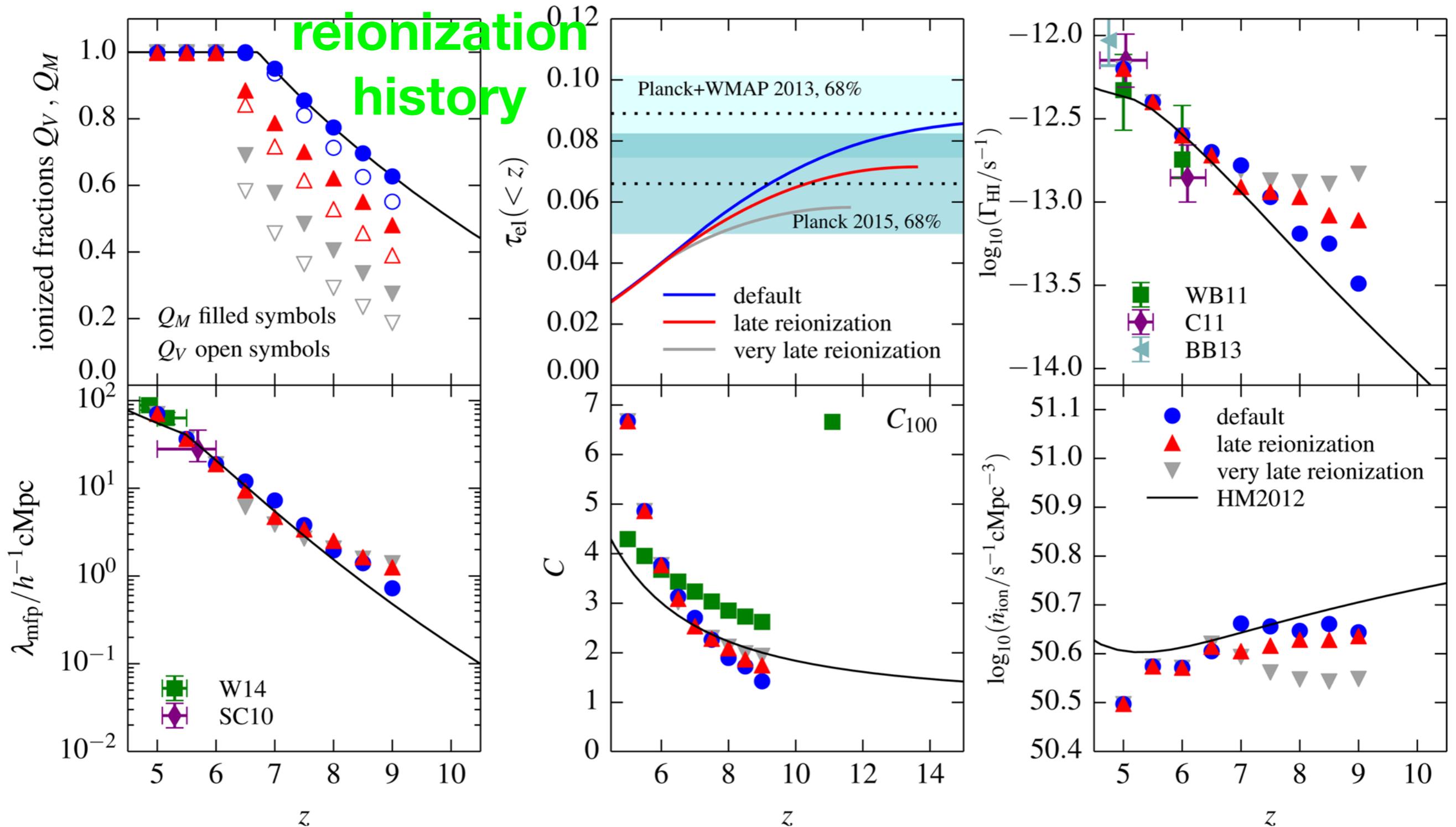


Iterative calibration of the photoionization rate



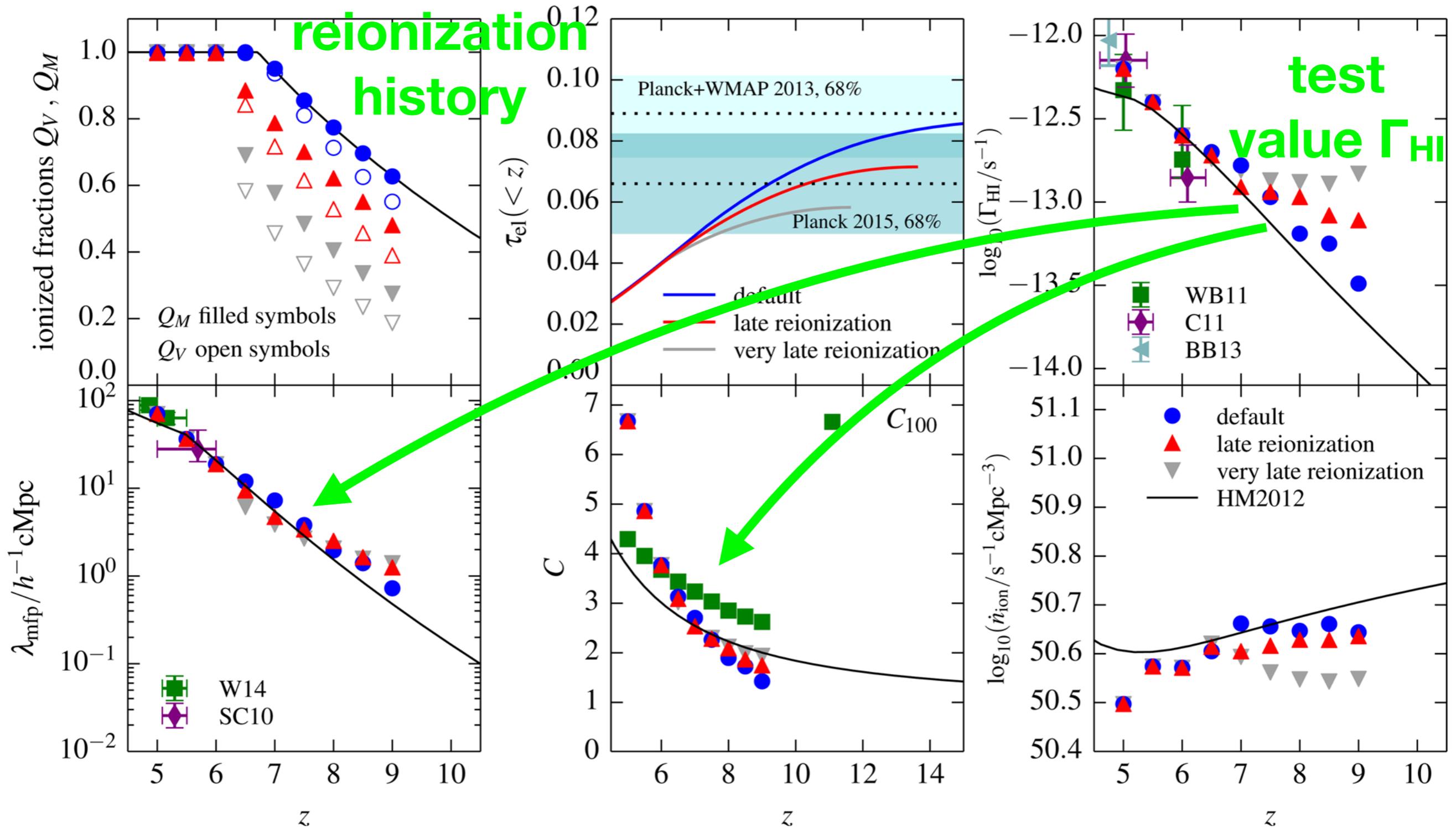
Iterative calibration of the photoionization rate

choose

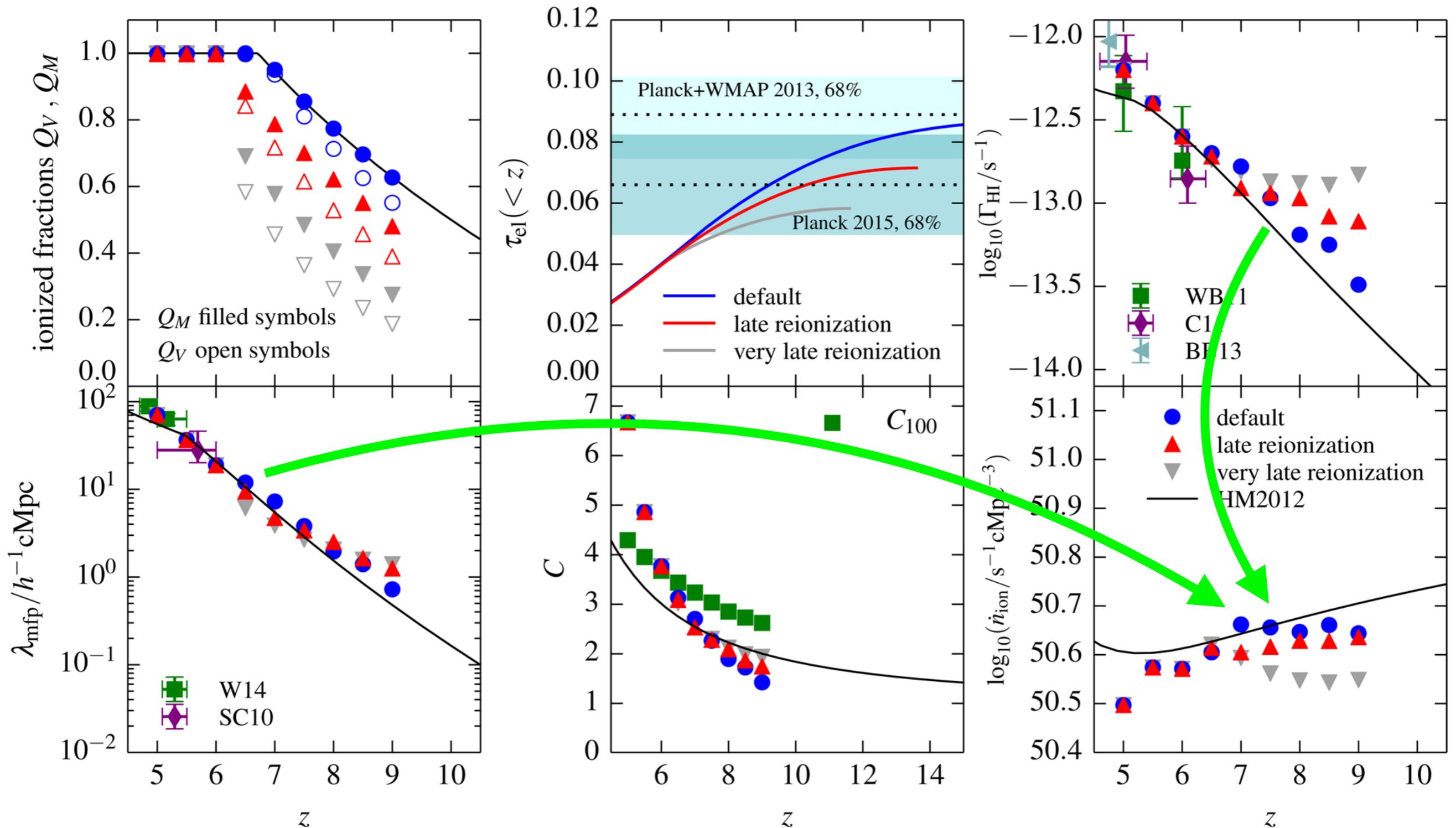


Iterative calibration of the photoionization rate

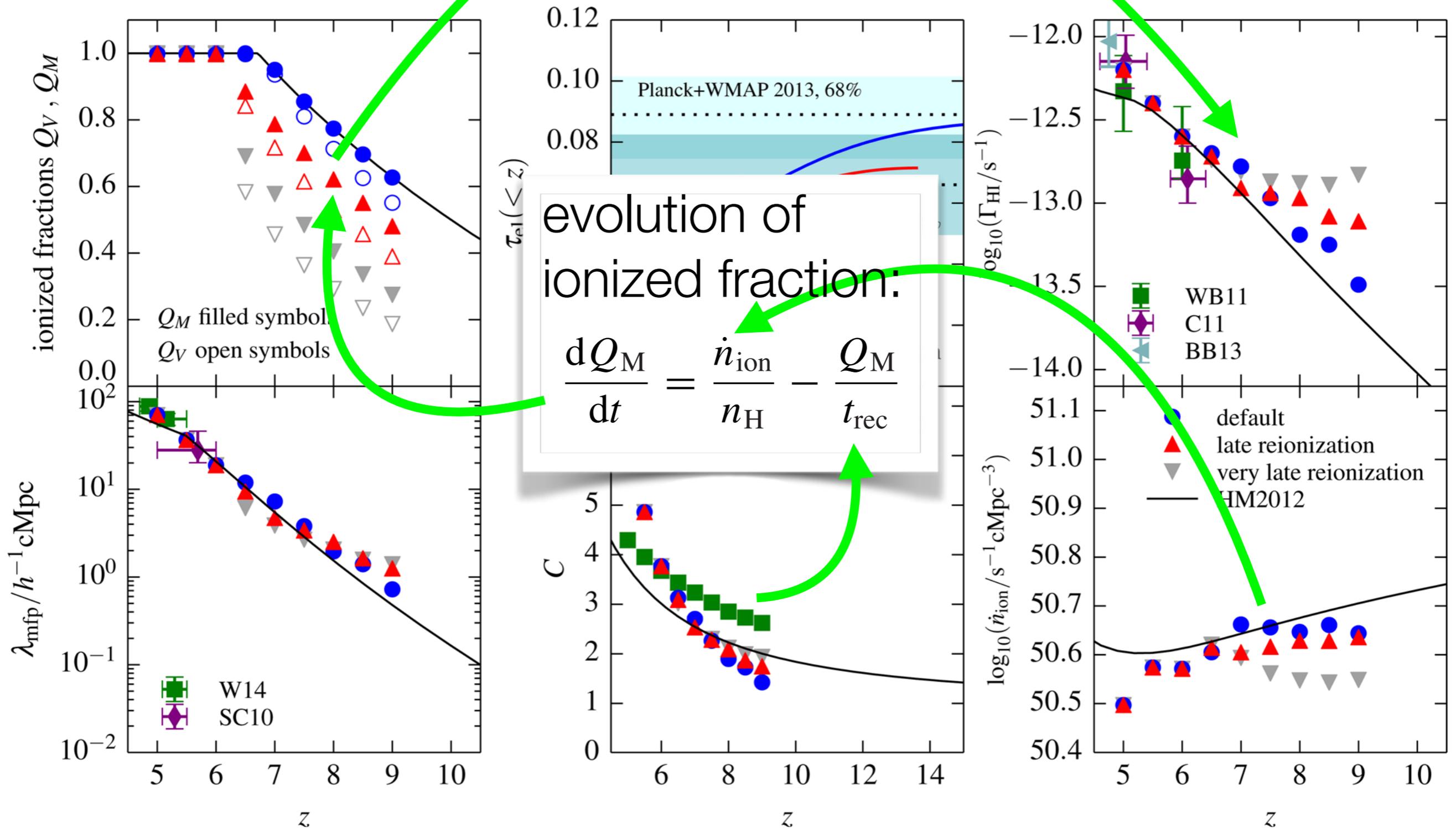
choose



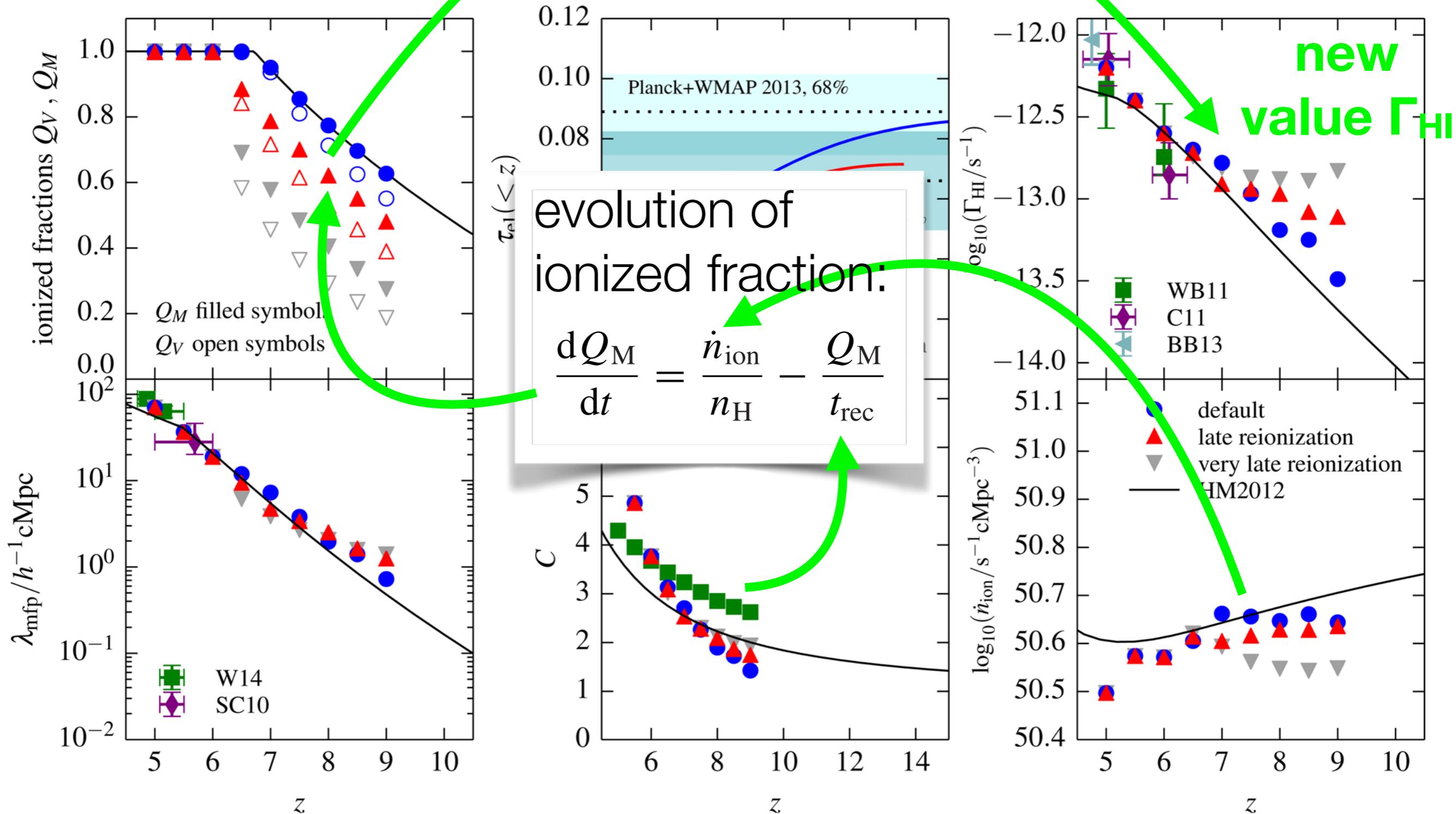
Iterative calibration of the photoionization rate



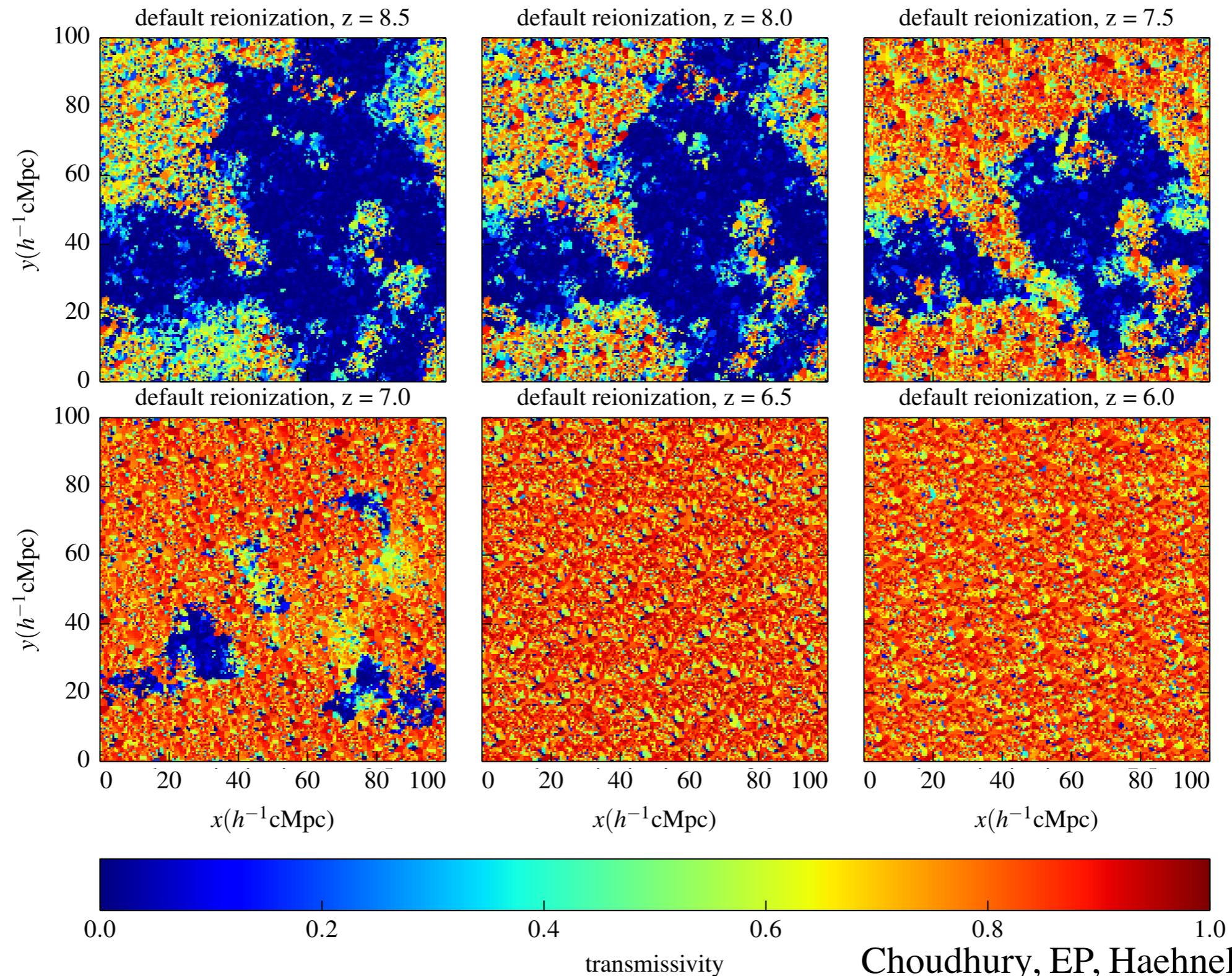
Iterative calibration of the photoionization rate



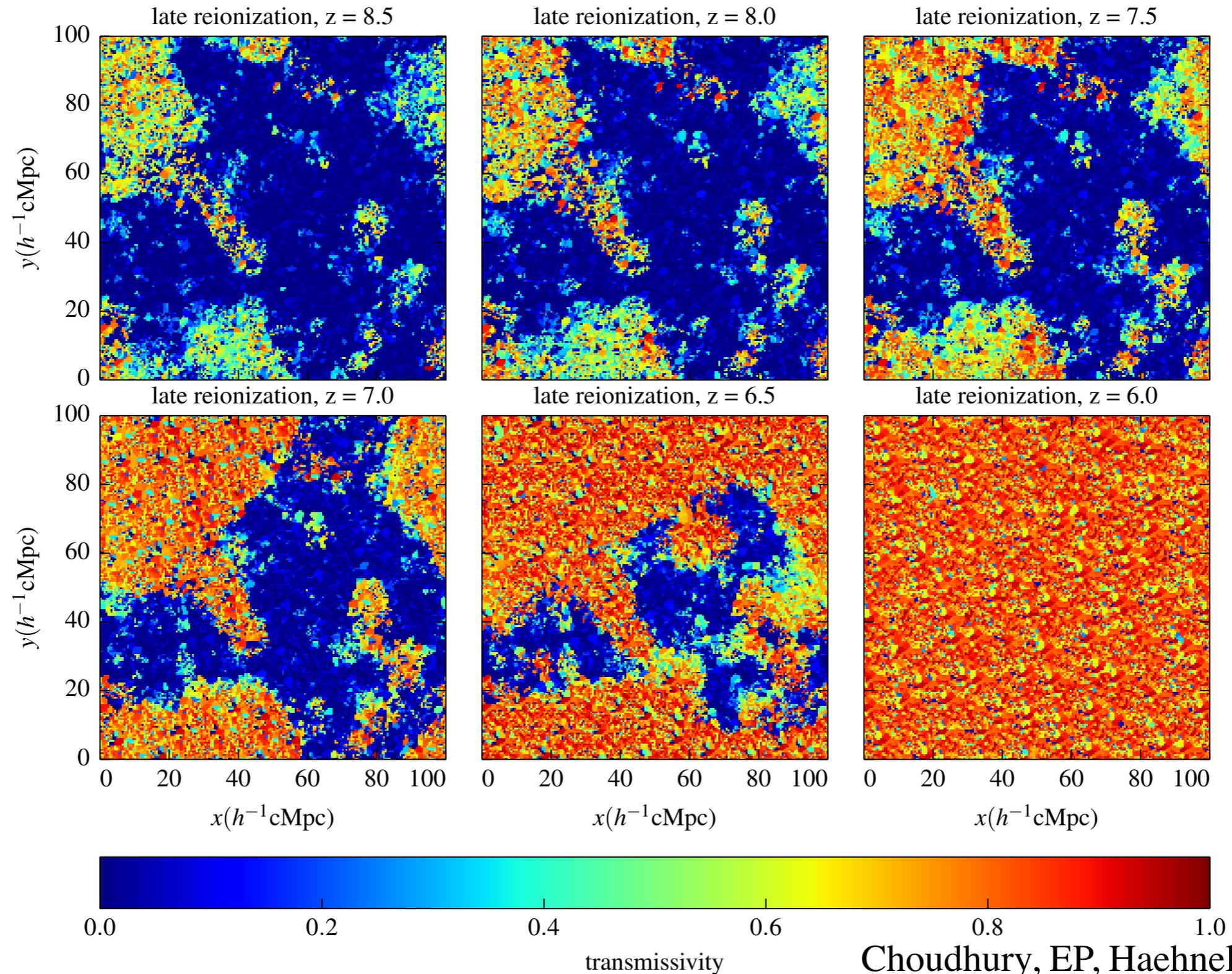
Iterative calibration of the photoionization rate



Transmissivity for Lyman- α emission lines



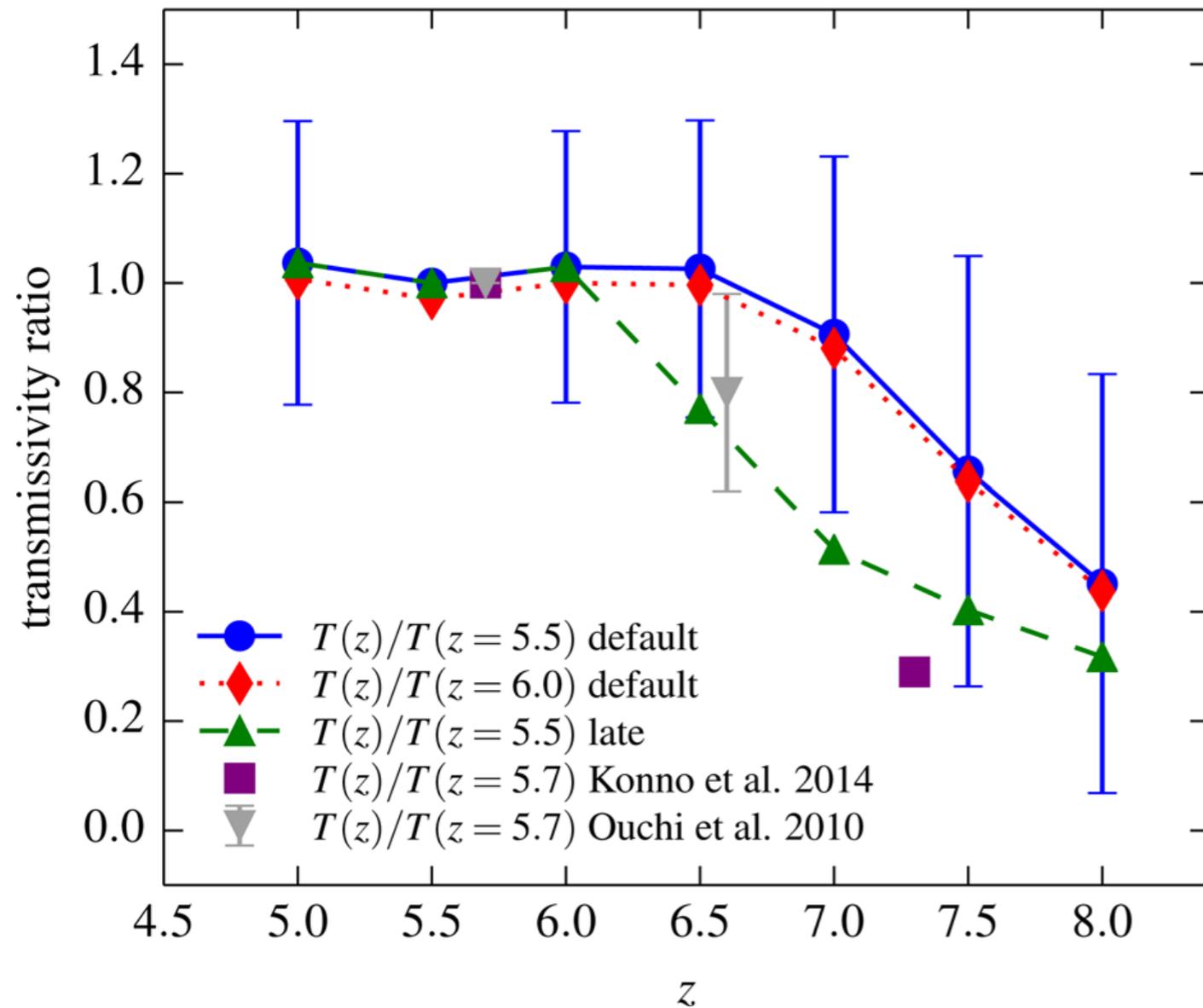
Transmissivity for Lyman- α emission lines



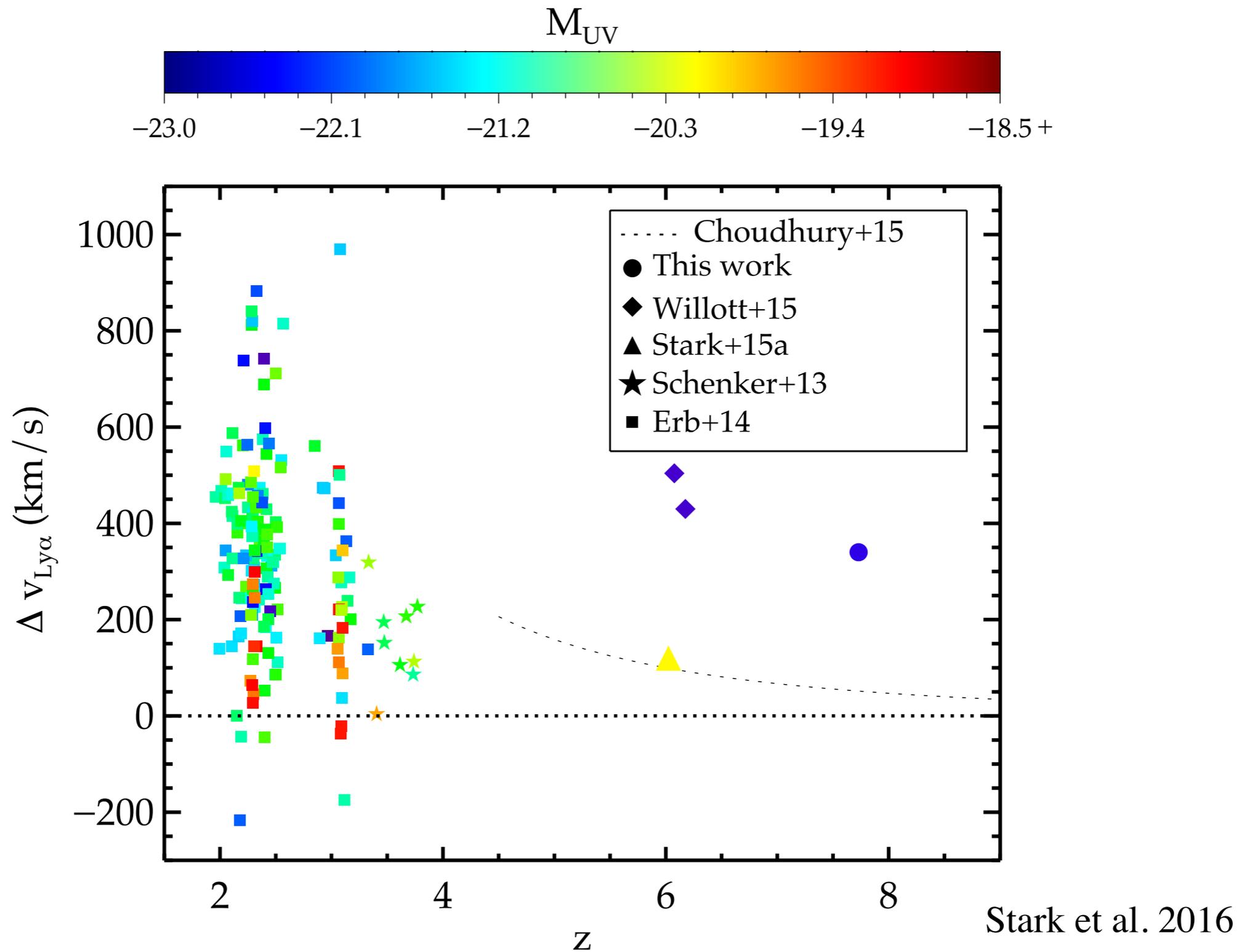
Redshift evolution of the transmissivity

constant velocity
shift

$$\Delta v_{\text{int}} = 100 \text{ km s}^{-1}$$



Intrinsic velocity shift vs redshift and luminosity



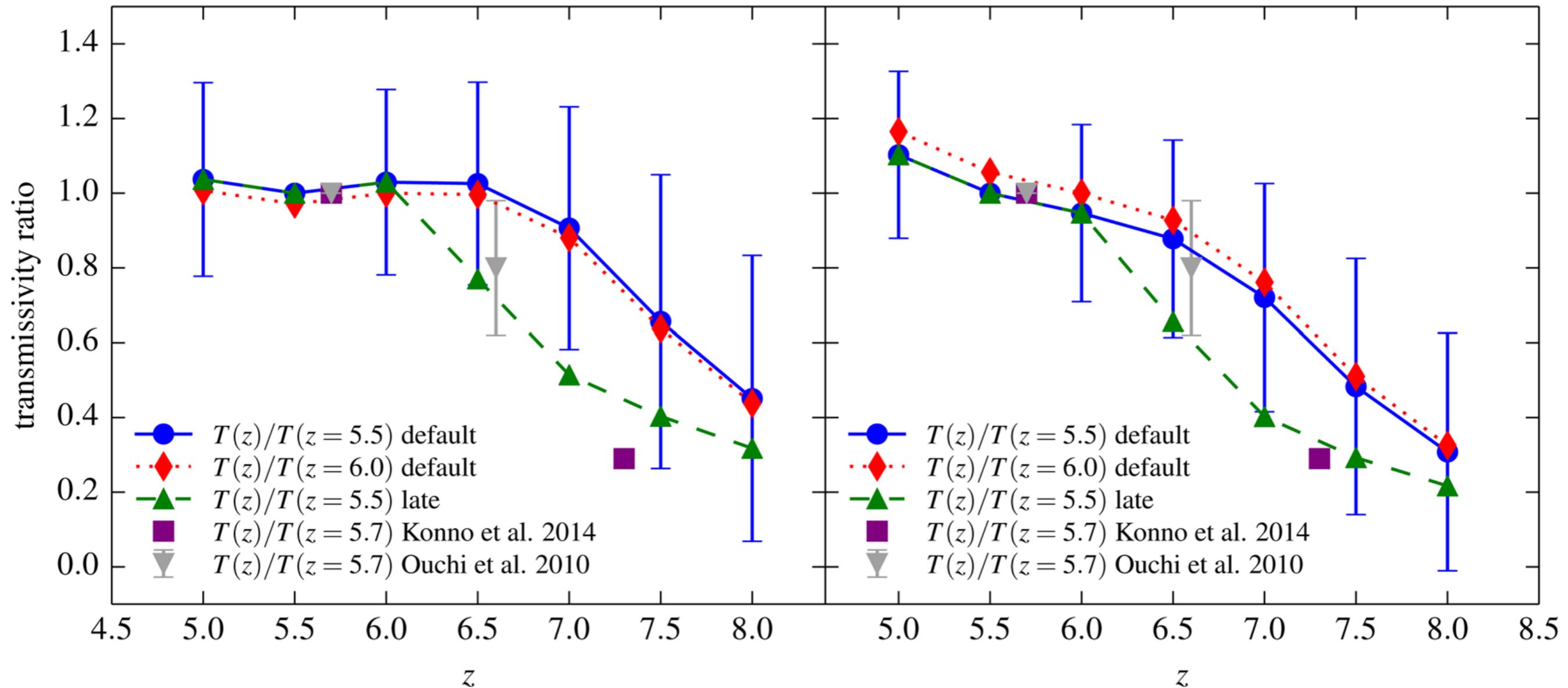
Redshift evolution of the transmissivity

constant velocity
shift

$$\Delta v_{\text{int}} = 100 \text{ km s}^{-1}$$

evolving velocity
shift

$$\Delta v_{\text{int}} = 100[(1+z)/7]^{-3} \text{ km s}^{-1}$$

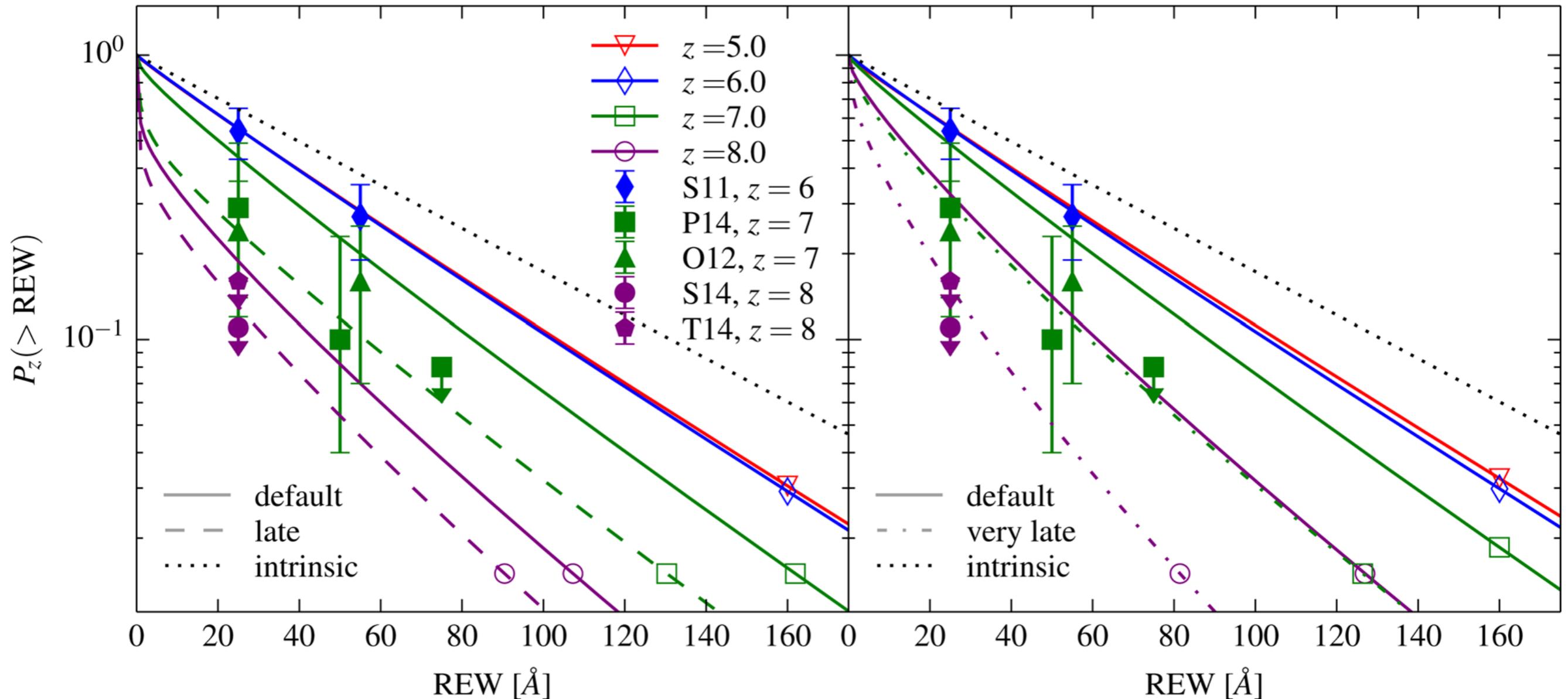


Cumulative Lyman- α equivalent width distribution

evolving velocity shift

uncorrelated LAE positions
 $\Delta v_{\text{int}} = 100[(1+z)/7]^{-3} \text{ km s}^{-1}$

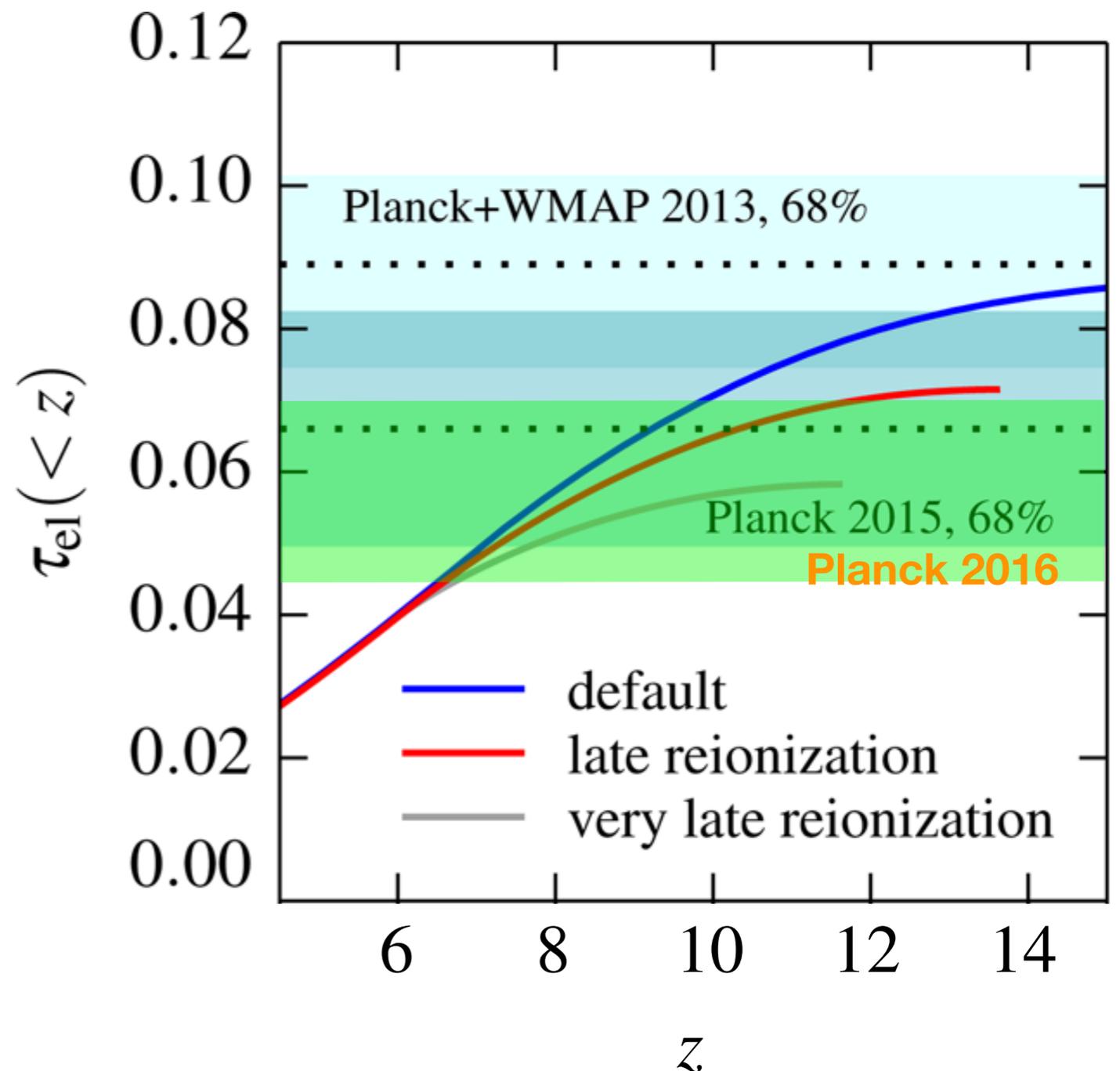
strongly correlated LAE positions
 $\Delta v_{\text{int}} = 100[(1+z)/7]^{-3} \text{ km s}^{-1}$



Consistency with latest Planck constraints

- optical depth in Planck 2015/16 came down
- consistent with what we need for the Lyman- α emitters
 - > reionization ends around $z \sim 6$ and is not too extended

Thomson scattering optical depth



Summary

- **Lyman- α emitters:**
 - favour a late and not too extended reionization history (finishing at $z \sim 6$)
 - evolution of intrinsic velocity offsets may be important
- **CMB:**
 - Planck 2015/16 find lower optical depths \rightarrow late reionization in agreement with LAEs