

$$73.8 \pm 2.1 \text{ km/s Mpc}$$

1. CEPHEIDS

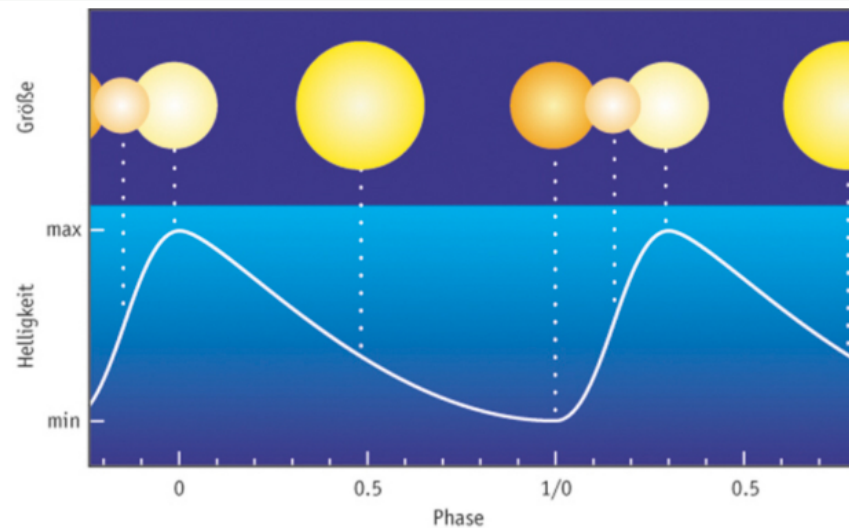
1.1 DISCOVERY

- 1784, Delta Cepheid in Cepheus
- 1912 discovery of another Cepheid in the magelanic cloud
- their Luminosity is proportional to their period



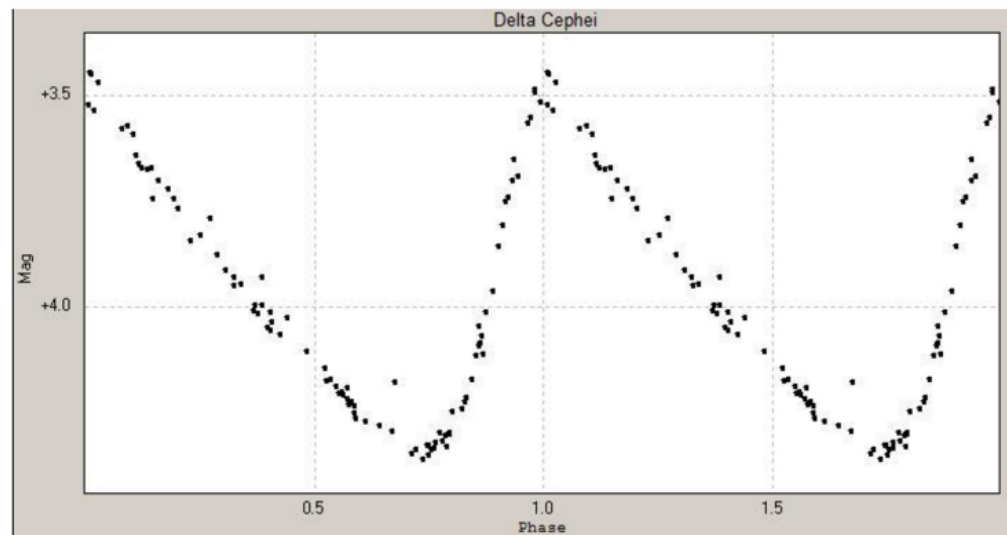
1.2 STRUCTURE

- determined by opacity of matter
- matter very opaque
- matter nearly transparent
- oscillate between two states:
 - Compact state
 - expanded state



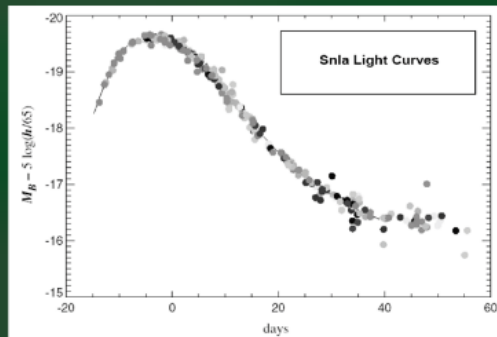
1.3 WHAT ARE THEY USED FOR?

- Distance determination through their period
- Cosmological ruler
- Determine the Hubble constant
- With the HST distance calculation to 20 Mpc



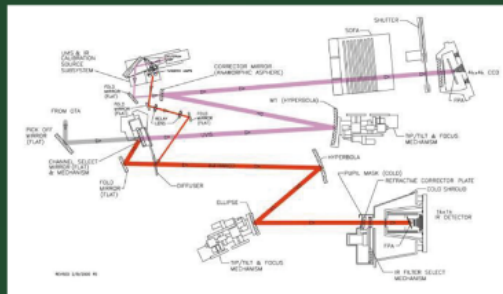
2. MEASURING OF THE HUBBLE CONSTANCE WITH CEPHEIDS

2.1 SUPERNOVA Ia

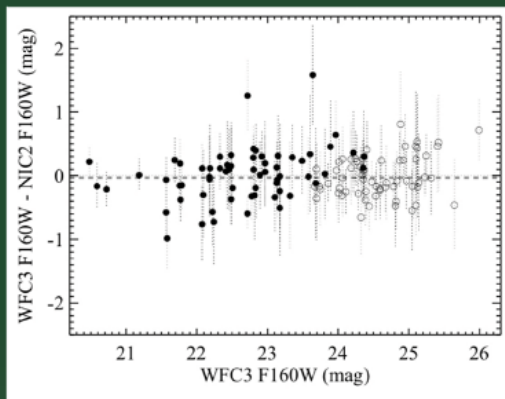


- Light-curves always the same
- cosmological candle
- use for Cross calibration of the Wide Field Camera 3

2.2 WIDE FIELD CAMERA 3, WFC3



- Observation from 200–1700 nm
- Two channels: IR red, UVIS Pink
- mechanism sends incoming light into desired channel

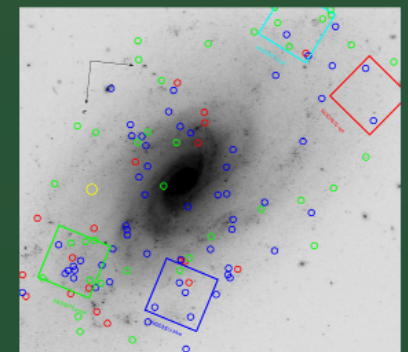
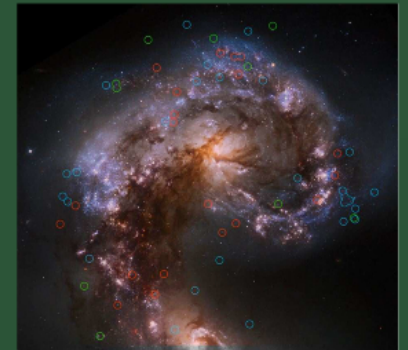
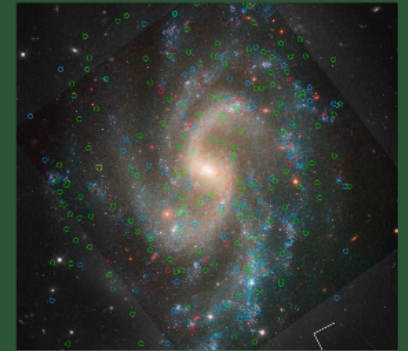


- it discovered 730 new Cepheids in 8 Host galaxies
- near IR with single photometric systems
- calibration with the same zeropoint
- Signal-to-Noise Ratio in a quarter of exposure time

2.3 DETERMINATION OF THE HUBBLE CONSTANCE

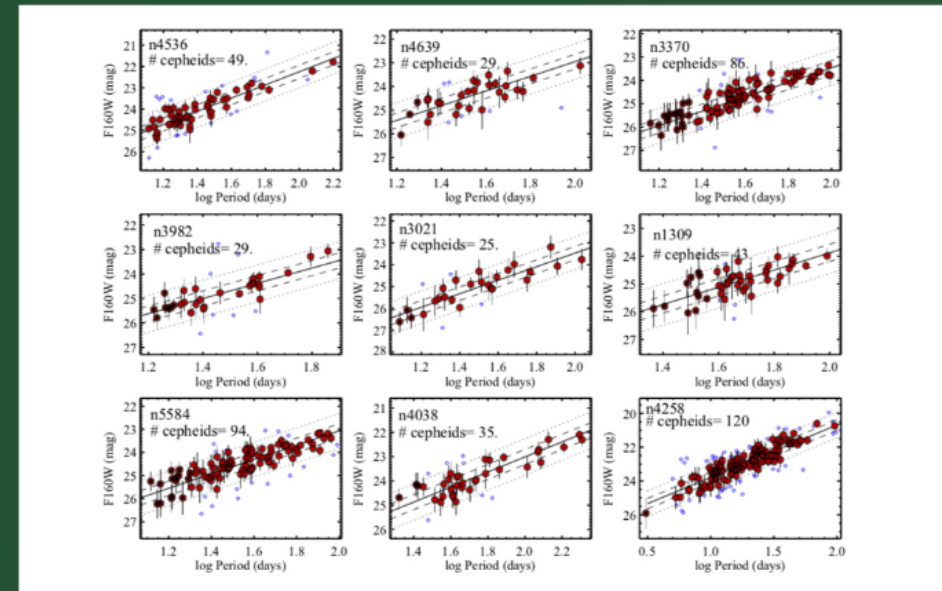
SYSTEMATICS

- 7% systematic error because of different zeropoints
- measurements of the flux of Cepheids
- use of Cepheids with similar metallicity and period
- we can measure the Hubble constant with a precision of 4,7%
- Cepheids in the NGC 5584, NGC 4038 and NGC 3370
- distance anchor NGC 4258



SOFTWARE CALCULATION

- automatically calculation of the WFC3 data
- Position uncertainty of the Cepheids is under 2,4 ms
- the P-L relation has outliers
- after outliers reduction: 484 Objekts Left



CALCULATING H_0

With Cepheids in all galaxies:

$$m_{W,i,j} = (\mu_{0,i} - \mu_{0,4258}) + z_{P_{W,4258}} + b_W \log P_{i,j} + z_W \Delta \log [O/H]_{i,j}$$

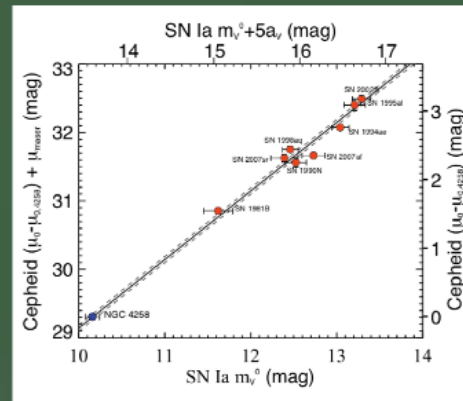
$$m_{W,ij} = m_{H,ij} - R(m_{V,ij} - m_{I,ij}),$$

$$R \equiv A_H / (A_V - A_I)$$

$$m_{v,i}^0 = (\mu_{0,i} - \mu_{0,4258}) + m_{v,4258}^0.$$



inclusion of relationship
between host Masses
and calibrated SN Ia





$$\log H_0 = \frac{(m_{V,4258}^0 - \mu_{0,4258}) + 5a_V + 25}{5}$$





$73.8 \pm 2.1 \text{ km/sMpc}$
uncertainty 50% Lower

Improvements:

- distance uncertainty to host at 3%
- Independent calibration of the First Rung of the distance ladder
- High signal-to-noise ratio measurements of the Milky Way parallaxes

Measuring the Hubble constant
with just Milky Way Cepheids

$$m_{W,i,j} = \mu_{0,i} + M_{W,i} + b_W \log P_{i,j} + Z_W \Delta \log [O/H]_{i,j}$$

$$M_{W,i,j} = M_{W,i} + b_W \log P_{i,j} + Z_W \Delta \log [O/H]_{i,j}$$

$$m_{v,i}^0 = \mu_{0,i} - M_v^0$$

$$\log H_0 = \frac{M_v^0 + 5a_v + 25}{5}.$$





$$75.7 \pm 2.6 \text{ km/sMpc}$$





with using the parallaxes
and the distance to NGC 4258
we get

$$H_0 = 74,5 \pm 2,3 \text{ km/sMpc}$$



and using an improved H-band
and the distance modulus of the Cepheids
we get

$$H_0 = 71,3 \pm 3,8 \text{ km/sMpc}$$



with all parameters combined
we get



$739.8 \pm 29.1 \text{ km/sMpc}$
uncertainty: 2,9% Lower

2.4 ERRORS AND IMPACTS ON THE HUBBLE CONSTANCE

CHEMICAL ABUNDANCE

- Calibration of Log [O/H]
- Te-scale
- reduces Log [O/H]
- that means the mean apparent metallicity: $12 + \text{Log [O/H]}$
- using just the Milky Way Cepheids, H_0 is increased by 2,0 km/sMpc
- metallicity correction:
-0,10 \pm 0,09 mag/dex

SUPERNOVE SYSTEMATICS

- Variants of Supernove measurement
- use SALT-II light-curve fitter
- with SALT-II we get
 $H_0 = 73,8 \pm 2,4 \text{ km/sMpc}$
with uncertainty of 3,3%

3. EXAMPLE FROM THE ASTROLAB

- period: $P = 50,4542$ d
- mean mag = $23,6 \pm 0,23$ mag
- $M = -2,760 * \text{Log}(P-1,0) - 4,160 = -12,22$ mag
- Distance modulus: $D = m - M = 35,82$ pc

4. FUTURE FOR THE HUBBLE CONSTANCE

- with the launch of JWST, hope to reduce the error of the first rung
- better cross calibration, with more Supernove, because a bigger observation sample
- not just observe 30 Mpc but 50 Mpc observations
- With this better value we could interpret a new species for neutrinos
- also we have a more exact value for the dark energy parameter $w = -1,08 \pm 0,10$



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